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# TIMES OF MINIMA OF ECLIPSING CATACLYSMIC VARIABLES

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#### Observatory and telescope:

 $0.6\mathrm{m}$  Ritchey-Chrétien (f/8) telescope (IST60) at Ulupinar Astrophysical Observatory, Canakkale.

Detector:	Apogee Alta U42 CCD camera, $2048 \times 2048$ pixels with
	a read-out noise of $10e^-$ RMS; SBIG STL-1001E CCD
	camera, $1024 \times 1024$ pixels with a read-out noise of $14.8e^-$
	RMS.

# Method of data reduction:

Reduction of the CCD frames was made in the usual way using  $IRAF^1$  package.

#### Method of minimum determination:

The minima times were computed with Kwee & Van Woerden (1956) method.

Times of minima:						
Star name	Time of min.	Error	Type	Filter	Rem.	
	HJD 2400000+					
BH Lyn	56195.5590	0.001	Ι	White-light	C2	
TT Tri	56195.2590	0.002	Ι	White-light	C2	
	56195.3990	0.003	Ι	White-light	C2	
	56570.3250	0.002	Ι	White-light	C1	
HS0455 + 8315	56193.3580	0.001	Ι	White-light	C2	
	56193.5070	0.001	Ι	White-light	C2	
	56571.2650	0.005	Ι	White-light	C1	
	56571.4140	0.005	Ι	White-light	C1	
PX And	56158.3920	0.002	Ι	White-light	C1	
	56194.5390	0.003	Ι	White-light	C2	
	56570.5200	0.001	Ι	White-light	C1	
V1315 Aql	56159.2930	0.001	Ι	White-light	C1	
	56159.4330	0.001	Ι	White-light	C1	

 $^{1}$ IRAF is distributed by the National Optical Astronomical Observatories, operated by the Association of the Universities for Research in Astronomy, inc., under cooperative agreement with the National Science Foundation

## Explanation of the remarks in the table:

C1 and C2 refer to the CCD cameras Apogee Alta U42 and SBIG STL-1001E, respectively.

### **Remarks:**

These objects were observed in the framework of a project which is carried out at Istanbul University to follow period changes in cataclysmic variables. HS0455+8315 is a SW Sex-type cataclysmic variable which was identified by the Hamburg Quasar Survey (Hagen et al. 1995).

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References:

Hagen, H. J. et al., 1995, A&AS, 111, 195

Kwee, K., van Woerden, H., 1956, Bulletin of the Astronomical Institutes of the Netherlands, 12, 327