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**SEARCH FOR RAPID OSCILLATIONS AMONG
SEVEN NORTHERN CP STARS**

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Up to now, there are only about 45 rapidly oscillating Ap (roAp) stars known. They are located within an area of pulsational instability in the Hertzsprung-Russell diagram, at the main sequence, ranging in effective temperature from about 6600 K to 8500 K.

Photometric investigations for these stars show a period range of five to twenty five minutes, which is consistent with acoustic (p-mode) pulsations of low degree and high radial overtone (Kurtz et al. 2011).

The driving mechanism of their oscillation modes is still a matter of debate. But the most probable explanation seems to be the “classical” κ -mechanism operating in the hydrogen ionisation zone. Many physical processes could play a role in this context: for example the coupling with the magnetic field, the ability of the latter to freeze convection and allow the stratification of chemical elements (Balmforth et al. 2001).

Paunzen et al. (2012) published a list of roAp candidates based on their location in the Hertzsprung-Russell diagram which served us as a primary source. In addition, we selected targets from the catalogue of Ap, HgMn and Am stars by Renson & Manfroid (2009). Since most of the known roAp stars are classified as SrCrEu, we searched for such objects in the aforementioned reference. The fundamental parameters of our targets listed in Table 1 were calibrated as described in Paunzen et al. (2012). Luminosities are only derived if Hipparcos parallaxes were available.

The observations were done at the 0.75m APT at Fairborn Observatory (Arizona) either in Strömgren v or Johnson B depending on the brightness of the target. The integration time was set to 20 seconds with apertures of 30 and 45 arcseconds, respectively. No comparison stars were used because a good sampling of the possible periods are needed. A polynomial function was fitted to the sky subtracted data to account for the extinction and atmospheric variations. In Table 2 the observation log is given. Because all data have

an arbitrary zero point, the measurements for each individual star were scaled and then merged.

The final light curves were examined in more detail using the Period04 program (Lenz & Breger 2005), which performs a discrete Fourier transformation. Significant peaks with periods of more than one hour were subtracted. In Figure 1, the Fourier spectra of all light curves are shown. There is no peak, which exceeds 1.3 mmag. For BD +35 3616, we notice a quite prominent peak at about 5 mHz which corresponds to about 200 seconds. Such a high frequency was never detected in any roAp star, so far. However, we are not able find to a possible explanation for this peak.

Table 1: Fundamental data of our target stars.

Star ID	V	Spec	T_{eff} [K]	$\log(L/L_{\odot})$
BD +25 4371	9.45	F0 Sr Cr Eu	8040(160)	–
BD +35 3616	9.47	F0 Sr Eu	6830(410)	–
HD 62140	6.47	A8 Sr Eu	7800(140)	1.30(7)
HD 94427	7.36	A5 Sr Eu Cr	7500(250)	1.29(10)
HD 96707	6.08	A8 Sr	7820(230)	1.55(6)
HD 98088	6.11	A8 Sr Cr Eu	8030(120)	1.66(12)
HD 115606	8.55	A2 Sr	7880(250)	1.20(21)

Table 2: Observation log.

Star ID (BD ...)	HJD(start) 2456200+	HJD(end) 2456200+	N	Star ID (HD ...)	HJD(start) 2456200+	HJD(end) 2456200+	N
+25 4371	61.5687	61.6422	254	62140	64.0183	64.0465	103
	62.5603	62.6340	256		66.0085	66.0244	62
	64.5695	64.5695	224		86.9491	86.9840	69
+35 3616	21.6505	21.6682	48	94427	73.9817	74.0476	224
	23.5663	23.6548	242		74.9755	75.0449	235
	25.6500	25.6570	260		96707	69.9945	70.0507
	27.5633	27.6461	232	71.9850		72.0516	224
	28.5686	28.6459	216	98088	90.9760	90.9946	66
	29.5620	29.6451	232		82.0011	82.0535	138
	30.5615	30.6446	234	115606	82.9552	83.0515	309
	31.5763	31.6405	175		86.9870	87.0141	93
	32.5628	32.6338	138				
	33.5605	33.6331	205				
	34.5589	34.6325	204				

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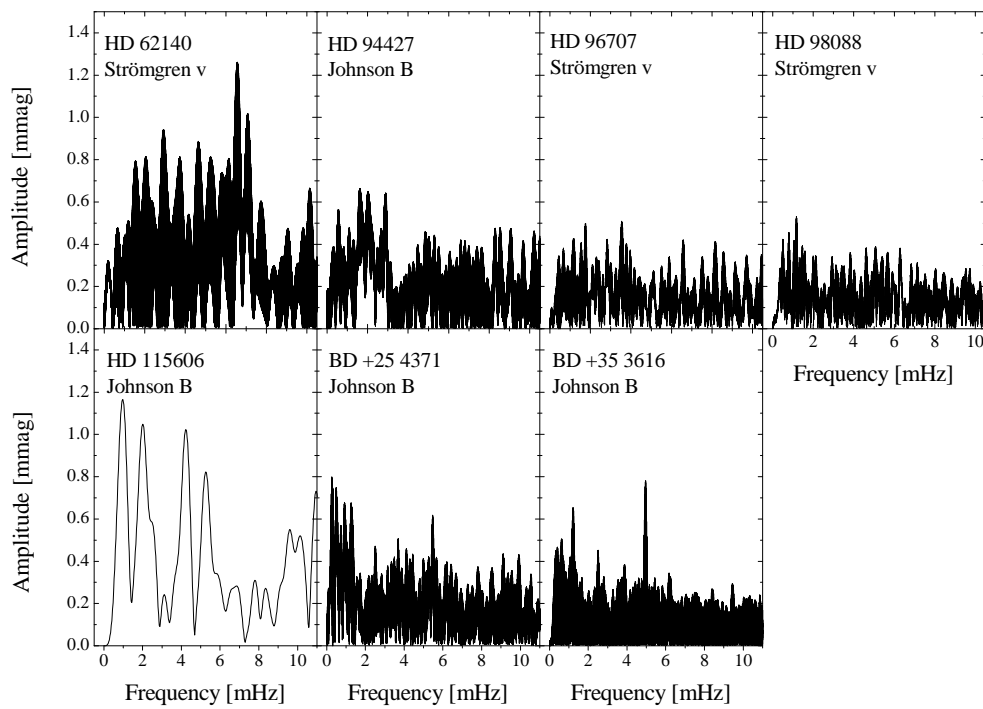


Figure 1. Fourier spectra of the target star light curves.

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