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NEW OUTBURST OF V1118 Ori (2007-2008)

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| | |
|---|-----------------------------------|
| Name of the object: | V1118Ori |
| Equatorial coordinates: | Equinox: |
| R.A.= 5 ^h 34 ^m 44 ^s .2 DEC.= -5°33'40" | |
| 2000 | |
| Observatory and telescope: | |
| Private obs., Sevilla(Spain) with Schmidt-Cassegrain tel. | |
| Detector: | CCD Camera |
| Filter(s): | V |
| Date(s) of the observation(s): | |
| 2007.02.09 – 2008.21.02 | |
| Comparison star(s): | Parenago 1492,1518,1540,1600,1641 |
| Availability of the data: | |
| Available at the IBVS website (5829-t1.txt) | |
| Type of variability: | EXor |
| Remarks: | |
| Since 1983, the discovery, V1118 Ori became known as an EXor or Subfuor (Parsamian and Gasparian, 1987; Herbig, 1990). We have information concerning its outbursts the periods 1983-84 (Kosai, 1983; Hurst et al., 1984; Parsamian and Gasparian, 1987), 1988-90 (Parsamian et al., 1993; Parsamian et al., 1996), 1992-94 (Garcia Garcia, Mampaso and Parsamian, 1995; Parsamian et al., 2002), 1996-98 (Hayakawa et al., 1998; Garcia Garcia and Parsamian, 2000), 2004-06 (Waagen et al., 2005; Williams et al., 2005; Garcia Garcia et al., 2006). New observations show, that V1118 Ori started brightening at 12.10.2007 until reached its maximum of V=15 ^m 3, then decreased. Some fluctuations of the brightness are observed. The last observation at 10.02.2008 is V=16 ^m 11. | |

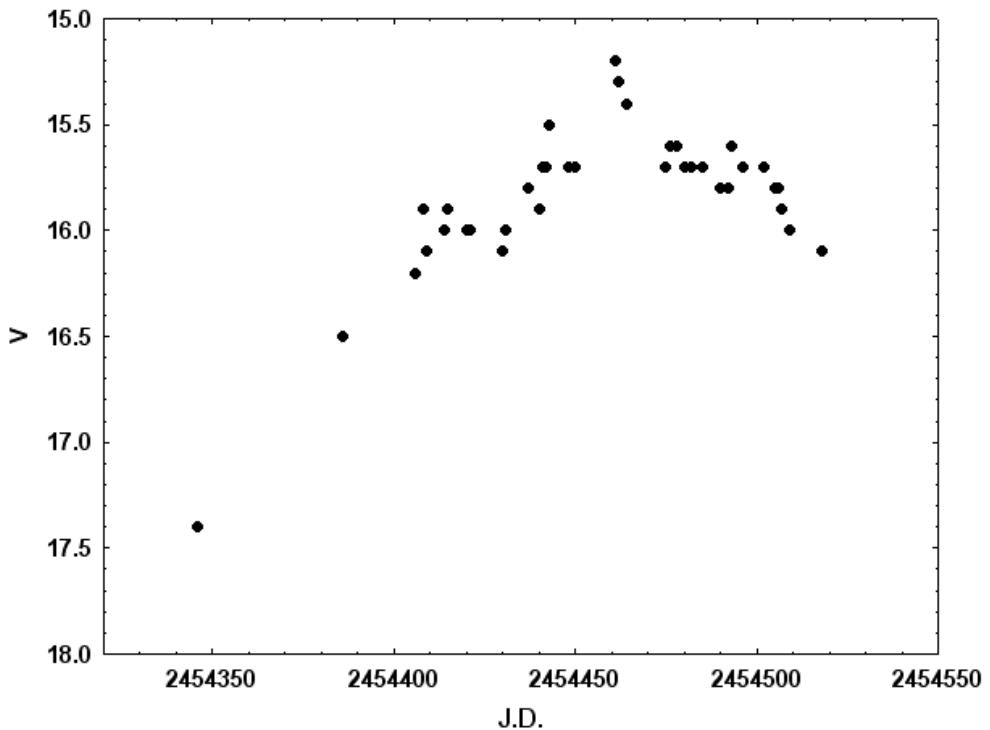


Figure 1.

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