# NEW TIMES OF MINIMA OF ECLIPSING BINARY SYSTEMS 

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| Observatory and telescope: |
| :--- |
| $50-\mathrm{cm} f / 8.4$ Ritchey-Chrétien telescope (Ba50) of the Baja Astronomical Obser- |
| vatory (Hungary) |
| $50-\mathrm{cm} f / 6$ modified Cassegrain telescope (Baja Astronomical Robotic Telescope - |
| BART1) of the Baja Astronomical Observatory (Hungary) |
| $50-\mathrm{cm} f / 15$ Cassegrain telescope (Pi50) of the Konkoly Observatory at Piszkéstető |
| Mountain Station (Hungary) |
| 25, and 40-cm Newton telescopes (Be25, Be40, respectively; Belgium) |
| 30-cm Cassegrain telescope of Setec Observatory, Kansas (Se30) |


| Detector: | $512 \times 512$ Apogee AP-7 CCD camera (Ba50) |
| :--- | :--- |
|  | $765 \times 510$ SBIG ST-7 CCD camera (Ba50ST7) |
| $4096 \times 4096$ Apogee Alta U16 CCD camera (BART1) |  |
|  | cooled UBVRI Photometer (Pi50) |
|  | $2184 \times 1472$ SBIG ST10XME with filterwheel (filters |
| Bessell specifications) (Bexx) |  |
|  | SBIG ST8 with filterwheel (filters Bessell specifications) |
| (Se30) |  |

## Method of data reduction:

Reduction of Baja CCD frames was made with a customly developed IRAF ${ }^{\dagger}$ package, while the others were reduced by Mira-AP (6) and (7)*softwares.

[^0]Method of minimum determination:
The minima times were computed with parabolic fitting, and in some cases with linearized Pogson-method or Kwee-van Woerden method (Kwee \& van Woerden, 1956).

| Times of minima: |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Star name | Time of min. HJD 2400000+ | Error | Type | Filter | Rem. |
| XZ And | 54012.5539 | 2 | I | V | Bor/BART1 |
| AB And | 53936.4859 | 1 | I | $V$ | Bor/Ba50 |
| EP And | 54048.3612 | 1 | II | $V$ | Heg/BART1 |
|  | 54048.5641 | 1 | I | $V$ | Heg/BART1 |
| OO Aql | 53613.4327 | 2 | II | V | Bor/Ba50 |
|  | 53881.5279 | 5 | II | $R$ | Bor/Ba50 |
| V889 Aql | 53255.392 | 1 | I | $B, V, R$ | Bor/Ba50 |
| SS Ari | 54056.4176 | 1 | I | $V$ | Heg/BART1 |
| CL Aur | 53675.4626 | 3 | II | $R$ | Bor/Ba50 |
| IM Aur | 53326.4270 | 2 | I | $V$ | Bor/Ba50 |
|  | 53447.411 | 1 | I | $V$ | Bor/Ba50 |
|  | 53790.4257 | 2 | I | $V, R$ | Reg+Bor/Pi50 |
|  | 54015.5599 | 1 | II | $V$ | Bor/BART1 |
|  | 54043.6266 | 2 | I | $V$ | Bor/BART1 |
| IU Aur | 52957.4095 | 12 | II | $B, V, R$ | Bír/Ba50ST7 |
|  | 53035.3063 | 15 | II | $V, R$ | $\mathrm{Heg} / \mathrm{Ba} 50$ |
|  | 53764.4187 | 3 | I | $B$ | Be40 |
|  | 53773.4739 | 3 | I | $R$ | Kis/Ba50 |
|  | 53780.7217 | 3 | I | $V$ | Se30 |
|  | 53789.7804 | 26 | I | $V$ | Se30 |
|  | 53800.6456 | 3 | I | $V$ | Se30 |
|  | 53803.3690 | 2 | II | $V$ | Bor/Ba50 |
|  | 53813.3244 | 11 | I | $V$ | Be25 |
|  | 54003.5350 | 14 | I | $V, R$ | Bor+Reg+Kov/Pi50 |
|  | 54043.3875 | 4 | I | $V$ | Bor/BART1 |
| TZ Boo | 53802.4937 | 2 | II | $V, R$ | Bor/Ba50 |
|  | 53802.6449 | 3 | I | $V, R$ | Bor/Ba50 |
|  | 53803.5348 | 2 | I | $V, R$ | Bor/Ba50 |
| Y Cam | 53824.5101 | 3 | I | $R$ | Kis/Ba50 |
|  | 54039.3840 | 6 | I | $V$ | Bor/BART1 |
| AS Cam | 53830.405 | 1 | II | $R$ | Kis/Ba50 |
| DN Cas | 54066.4437 | 4 | I | $V$ | Be40 |
| PV Cas | 53183.5042 | 3 | II | $V$ | Bor/Ba50 |
| VW Cep | 53608.4033 | 7 | II | $B, V, R$ | Bor/Ba50 |
|  | 53848.4473 | 2 | I | $V$ | Bor/Ba50 |
|  | 53848.5869 | 1 | II | $V$ | Bor/Ba50 |
|  | 53892.4210 | 9 | I | $B, V, R$ | Reg+Bor /Pi50 |
|  | 53947.385 | 1 | II | $V, R$ | Kov+Reg/Pi50 |
| XX Cep | 54004.4338 | 4 | I | $V, R$ | Bor+Kov+Reg/Pi50 |
|  | 54018.4576 | 2 | I | $V$ | Bor/BART1 |
| EK Cep | 53745.2544 | 19 | II | $V$ | Be25 |
| LS Del | 53937.530: | 3 | I | $B, V, R$ | Heg/Ba50 |
|  | 53938.4305 | 3 | II | $V$ | Bír/BART1 |
| DI Her | 53933.4810 | 4 | I | $V$ | Bor/Ba50 |
| HS Her | 53935.4277 | 4 | I | $V$ | Bor/Ba50 |
| V994 Her | 53206.365 | 2 | ? | $V, R$ | Bor/Ba50 |
| SW Lac ${ }^{\text {a }}$ | 53596.5127 | 1 | II | $R$ | Bor/Ba50 |
|  | 53596.5136 | 1 | II | $V$ | Bor/Ba50 |
|  | 54015.3755 | 1 | II | $V$ | Bor/BART1 |
| AR Lac | 54001.4618 | 8 | II | $B, V, R$ | Reg+Bor/Pi50 |
| AU Lac | 53745.2926 | 2 | I | $V$ | Be40 |


| Times of minima: |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Star name | Time of min. <br> HJD 2400000 + | Error | Type | Filter | Rem. |
| UV Leo | 53459.3746 | 3 | II | $R$ | Bor/Ba50 |
|  | 53797.5236 | 3 | I | $V, R$ | Bor/Pi50 |
|  | 53828.4280 | 2 | II | $R$ | Bor/Ba50 |
| U Peg | 54000.544 | 1 | II | $V, R$ | Bor/Pi50 |
| AG Per | 54034.429 | 1 | I | $V, R$ | Kov+Bor+Reg/Pi50 |
|  | 54039.5241 | 5 | II | $V$ | Bor/BART1 |
| $\beta \mathrm{Per}^{\text {b }}$ | 54084.360 | 3 | II | $(V, R)+N$ | Bor+Reg/Pi50 |
| EQ Tau | 53802.3811 | 4 | I | $V, R$ | Bor/Ba50 |
|  | 53815.3525 | 2 | I | $R$ | Bor/Ba50 |
| TW UMa | 53813.4281 | 8 | I | - | Be25 |
| VV UMa | 53765.5233 | 1 | I | $V$ | Be40 |
| ZZ UMa | 53814.4329 | 1 | I | $V$ | Be25 |
| DW UMa | 53080.5071 | 1 | I | $R$ | Bor/Ba50 |
|  | 53080.6434 | 1 | I | $R$ | Bor/Ba50 |
|  | 53437.3241 | 1 | I | $V$ | Bor/Ba50 |
|  | 53443.4711 | 2 | I | $R$ | Bor/Ba50 |
|  | 53443.6082 | 2 | I | $R$ | Bor/Ba50 |
|  | 53451.3942 | 1 | I | $V$ | Bor/Ba50 |
|  | 53767.3656 | 2 | I | $R$ | Bor/Ba50 |
|  | 53815.4506 | 1 | I | $R$ | Bor/Ba50 |
|  | 53815.5869 | 1 | I | $R$ | Bor/Ba50 |
|  | 53822.4174 | 2 | I | $R$ | Bor/Ba50 |
|  | 53861.3504 | 1 | I | $R$ | Bor/Ba50 |
|  | 53861.4875 | 2 | I | $R$ | Bor/Ba50 |
| LP UMa | 53080.5012 | 3 | II | $R$ | Bor/Ba50 |
|  | 53443.5454 | 7 | I | $V, R$ | Bor/Ba50 |
|  | 53451.4473 | 4 | II | $V$ | Bor/Ba50 |
|  | 53767.393 | 2 | I | $R$ | Bor/Ba50 |
|  | 53815.5834 | 4 | II | $R$ | Bor/Ba50 |
|  | 53819.4545 | 2 | I | $V$ | Bor/Ba50 |
|  | 53822.4011 | 3 | II | $R$ | Bor/Ba50 |
|  | 53861.4435 | 4 | II | $R$ | Bor/Ba50 |

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Explanation of the remarks in the table:
Observer(s)/Instrument
a}\mathrm{ : SW Lac: On the night 53596 the discrepancy between the mid-eclipse time in V
and R band is supposed to be real.
}\mp@subsup{}{}{b}:\beta\mathrm{ Per: Due to the brightness of the system we had to use an additional neutral
filter (denoted by N).
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## Reference:

Kwee, K. K., van Woerden, H., 1956, Bull. Astron. Inst. Neth., 12, 327


[^0]:    $\dagger$ IRAF is distributed by the National Optical Astronomical Observatories, operated by the Association of the Universities for Research in Astronomy, inc., under cooperative agreement with the National Science Foundation
    *Mira software is produced by Mirametrics Inc.

