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TIMES OF MINIMA OF THE ECLIPSING BINARY SYSTEM EG CEPHEI

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Observatory and telescope:
Mount Laguna Observatory, 0.4-m and 0.6-m reflectors

Detector:

Photoelectric (see Remarks)

Method of data reduction:

Standard photoelectric differential photometry reduction

Method of minimum determination:

Kwee–van Woerden algorithm

Observed star(s):								
Star name	GCVS	Coordinates $(J2000)$		Comp./check				
	type	$\mathbf{R}\mathbf{A}$	Dec	$\operatorname{star}(\mathrm{s})$				
EG Cep	EB	$20^{h}15^{m}56.8$	$+76^{\circ}48'36''$	HD 193834, HD 194400				

Times of minima:									
Star name	Time of min.	Error	Type	Filter	Rem.				
	${ m HJD}~2400000+$								
EG Cep	2439004.7837	3	Р	UBV					
	2439137.6721	4	Р	UBV					
	2439290.7096	3	Р	UBV					
	2439292.8882	5	Р	UBV					
	2439297.7895	3	Р	UBV					
	2447732.8995	3	Р	uvb					
	2448067.8419	3	Р	uvby					
	2448121.7603	3	Р	uvby					

Remarks:

The first five times of minima were determined from data in an unpublished master's thesis (Cochran, 1967). Cochran's observations were made at Mount Laguna Observatory using a 0.4-meter reflecting telescope with a dry ice cooled 1P21 photomultiplier, the UBV system, and a charge-integrating photometer. His comparison and check stars were HD 193834 and BD+76°787 respectively. EG Cephei was also observed during 1989 and 1990 at Mount Laguna Observatory with the 24-inch Smith reflector. The 1989 observations were made using a photometer employing an EMI 6256 photomultiplier, while the 1990 observations were made with a Hammamatsu R943-02 tube, both were thermoelectrically cooled. This photometry was carried out in pulse-counting mode using the Strömgren uvby system. Comparison and check stars were HD 194400 (F8, V = 9.72) and HD 194130 (F2, V = 8.87) respectively.

Acknowledgements:

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Reference:

Cochran, G.V., 1967, Masters Thesis, San Diego State University

ERRATUM FOR IBVS 5607

The correct identifier for NSV 10478 is USNO 0900-12232367.

The Editors

ERRATUM FOR IBVS 5681

One of the eccentric eclipsers in IBVS 5681 is wrongly identified as GSC 3682-0837 =USNO-A2.0 1425-02073759 = 2MASS J01315922+5926474.

The eclipsing binary with a period of 6.1772 d is actually GSC 3682-0736 = UCAC250208296 = 2MASS J01215916+5833136 at $01^{h}21^{m}59.16 + 58^{\circ}33^{m}13.6$ (2000.0). The spectral type is B0.

P. Dubovsky, S. Otero