COMMISSIONS 27 AND 42 OF THE IAU INFORMATION BULLETIN ON VARIABLE STARS

Number 5379

Konkoly Observatory Budapest 24 February 2003 HU ISSN 0374 - 0676

FIRST BVR PHOTOMETRY OF TV URSAE MINORIS

DEĞİRMENCİ, Ö. L.¹; BOZKURT, Z.¹; YAKUT, K.¹; DEMİRCAN, O.²; ERDEM, A.²; ÇİÇEK, C.²; ÖZDEMÍR, S.²; BULUT, I.²; SOYDUGAN, F.²; SOYDUGAN, E.²; ESENOĞLU, H.³; HEGEDÜS, T.⁴; BORKOVITS, T.⁴; BÍRÓ, I. B.⁴

¹ Ege University Observatory, TR-35100 Bornova, İzmir, Turkey; e-mail: omerd@astronomy.sci.ege.edu.tr

² Canakkale Onsekiz Mart University, Faculty of Arts and Sciences, Department of Physics, TR-17100 Çanakkale, Turkey

³ İstanbul University Observatory, TR-34452 İstanbul, Turkey

⁴ Baja Astronomical Observatory of Bács-Kiskun County, Baja, Szegedi út, P.O. Box 766, H-6500, Hungary

Name of the object:				
	$4 = \text{HIP} \ 73474 = \text{HD} \ 133767$			
		·		
Equatorial coordinates:		Equinox:		
$\mathbf{R.A.} = 15^{h}00^{m}59.69$ I	$DEC. = +73^{\circ}03'11''_{.5}$	2000		
Observatory and tele	escope:			
	tory, 48-cm Cassegrain telescope			
S 1	n Ritchey-Chrètien telescope			
	J - F -			
Detector:	Hamamatsu, R 4457 (PMT)			
	SiTE 502B (Apogee AP-7 CCD)			
Filter(s):	Johnson B, V and R			
	unfiltered (Baja)			
Comparison star(s):	$BD + 73^{\circ}660 = SAO 8166 (Ege)$			
	BD + 73°655 = SAO 8100 (Ege) BD + 73°655 = SAO 8148 (Baja)			
		/		
Check star(s):	$BD + 73^{\circ}645 = HD \ 131493$			
Transformed to a sta	ndard system: No			
Availability of the da	ata:			
Upon request				
Type of variability:	EB			

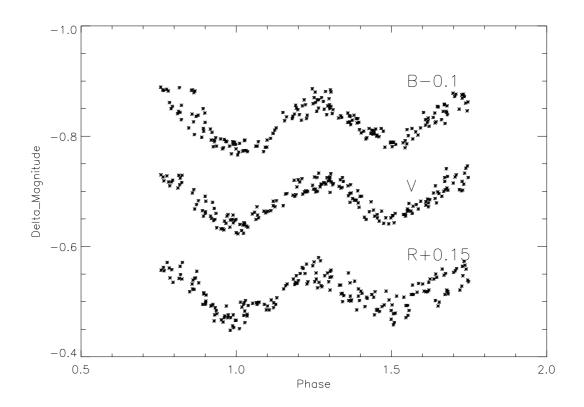


Figure 1. B, V and R light curves of TV UMi obtained at the Ege University Observatory.

Remarks:

In this paper we present B, V and R light curves of the eclipsing binary TV UMi. The variability of TV UMi was discovered by HIPPARCOS (ESA, 1997). The photometric observations of the system by HIPPARCOS show an EB type light curve (Kazarovets et al., 1999). The light elements of the system were given (ESA, 1997) as follows:

HJD Min. I = $2448500.2744 + 0.415546 \times E$.

The spectral type of the system is given as F8 (ESA, 1997). The first ground-based photometric observations were made over 9 nights during 2000 observing season at the Ege University Observatory and 3 nights at Baja Observatory as a part of the bilateral photometric campaign of the two institutes. The derived light curves for B, V, R colours are illustrated in Figure 1. The obtained minima times are given in Table 1. The O - C values are calculated on the base of our improved ephemeris (see below). Using the data given in Table 1 we derived the following, improved light elements and their errors:

HJD Min. I = $2448500.272(8) + 0.415550(1) \times E$.

The phases in Figure 1 are calculated using the above elements.

Times of a Star name	Time of min.	Error	Type	Filter	O-C	Rem.
	HJD 2400000+		-JP-		[day]	
TV UMi	48500.2744	_	Ι	_	0.0022	Hipparcos
	51684.421	2	II	B, V, R	-0.003	Ege
	51731.382	2	II	B, V, R	0.001	Ege
	51762.331	2	Ι	B, V, R	-0.009	$\overline{\text{Ege}}$
	51765.448	1	II	_	-0.008	Baja
	51766.487	1	Ι	_	-0.008	Baja
	51773.344	1	II	_	-0.008	Baja
	52091.448	1	Ι	B, V, R	-0.007	Ege
	52100.398	3	II	B, V, R	0.009	Ege

This work was partly supported by Hungarian National Grant OTKA T030743.

References:

ESA, 1997, The Hipparcos and Tycho Catalogues, SP–1200 Kazarovets, A. V., et al., 1999, IBVS, 4659