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**PHOTOELECTRIC MINIMA TIMES  
OF SOME ECLIPSING BINARIES**

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We present 24 photoelectric minima times of 5 selected eclipsing binary systems. All observations were obtained with the 30-cm Maksutov telescope at the Ankara University Observatory. Differential observations were secured by using an OPTEC SSP-5A photometer head which contains a side on R-1414 Hamamatsu photomultiplier. The filters used are in close accordance with the standard Johnson's *UBV* and reductions of the observations have been performed in the usual way (Hardie, 1962).

The moments of minima and their standard errors for each filter were calculated by using the well known method of Kwee & van Woerden (1956). All results are listed in Table 1 with the minimum types, filters and observers.

References:

- Hardie, R., 1962, *Astr. Tech.: Stars and Stellar Systems*, Vol. II, Univ. of Chicago Press, Chicago  
Kwee, K. K. & van Woerden, H., 1956, *Bull. Astron. Inst. Neth.*, **12**, 327

Table 1: Minima times of the observed systems

System	Min HJD 2400000 +	Error	Min type	Filter	Observer (*)
44i Boo	51851.26017	$\pm 0.00049$	I	<i>B</i>	Gr
	51851.26598	$\pm 0.00353$	I	<i>V</i>	Gr
WY Cnc	51958.37627	$\pm 0.00021$	I	<i>B</i>	Öz
	51958.37546	$\pm 0.00033$	I	<i>V</i>	Öz
SW Lac	51738.44838	$\pm 0.00056$	I	<i>U</i>	Ad
	51738.44924	$\pm 0.00093$	I	<i>B</i>	Ad
	51738.44911	$\pm 0.00069$	I	<i>V</i>	Ad
	51874.27251	$\pm 0.00045$	II	<i>U</i>	Ny
	51874.27339	$\pm 0.00043$	II	<i>B</i>	Ny
	51874.27334	$\pm 0.00036$	II	<i>V</i>	Ny
	51881.32895	$\pm 0.00067$	II	<i>U</i>	Tn
	51881.32785	$\pm 0.00018$	II	<i>B</i>	Tn
	51881.32784	$\pm 0.00022$	II	<i>V</i>	Tn
	51888.22210	$\pm 0.00028$	I	<i>U</i>	Dn
UV Lyn	51888.22195	$\pm 0.00020$	I	<i>B</i>	Dn
	51888.22185	$\pm 0.00021$	I	<i>V</i>	Dn
	51935.41371	$\pm 0.00028$	I	<i>U</i>	Gd
	51935.41595	$\pm 0.00032$	I	<i>B</i>	Gd
	51935.41661	$\pm 0.00021$	I	<i>V</i>	Gd
DI Peg	51949.51982	$\pm 0.00024$	I	<i>B</i>	Ak
	51949.52276	$\pm 0.00042$	I	<i>V</i>	Ak
	51035.40130	$\pm 0.00021$	I	<i>U</i>	Gr
DI Peg	51035.40058	$\pm 0.00028$	I	<i>B</i>	Gr
	51035.40044	$\pm 0.00021$	I	<i>V</i>	Gr

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