## COMMISSIONS 27 AND 42 OF THE IAU INFORMATION BULLETIN ON VARIABLE STARS

Number 5057

Konkoly Observatory Budapest 10 April 2001 HU ISSN 0374 - 0676

## V802 Aql IS AN ECLIPSING BINARY OF W UMa-TYPE

VAN CAUTEREN, P.<sup>1,2</sup>; WILS, P.<sup>2</sup>

V802 Aql (= GSC 5119-00948;  $\alpha_{2000} = 18^{\rm h}58^{\rm m}55^{\rm s}$ ;  $\delta_{2000} = -03^{\circ}1'.2$ ) was discovered by Bakos (1950) who classified it as an RR Lyrae star varying between magnitudes 13.7 and 14.2 with a period of 0.1338, based on 25 observed minima. Rodríguez et al. (2000) included the star in their latest catalogue of  $\delta$  Scuti stars.

We observed V802 Aql on two nights in 1996 and on four nights in 2000. We obtained respectively 130 and 347 data points, totalling 12.7 hours of photometry. We used a 0.40-m telescope. In 1996 this instrument was equipped with a Hisis24 CCD camera, in 2000 with a ST7 CCD camera. Both cameras have a Kodak KAF400 chip. No filter was used. Exposure times varied between 70 and 120 seconds, depending on sky conditions. The images were reduced with the aperture photometry procedure of the Mira AP software package<sup>†</sup>.

The brightness of the variable was measured with respect to the average of GSC 5119-01018 and GSC 5119-00575. The standard deviation of the differential magnitudes between both comparison stars was 0.006. The data were frequency-analysed with "Period", a period search program developed by one of us (P. Wils) based on the PDM method (Stellingwerf, 1978). The following times of minima were observed:

| Type   | Mag   | JD Hel.     |
|--------|-------|-------------|
| Min II | 0.295 | 2450300.434 |
| Min II | 0.325 | 2451780.457 |
| Min I  | 0.369 | 2451781.392 |
| Min I  | 0.349 | 2451782.463 |
| Min II | 0.288 | 2451784.471 |

Our results show that V802 Aql is an eclipsing binary of type W UMa with a period of  $0^{d}.2677 \pm 0^{d}.0003$ . Fig. 1 shows the phased light curve with two unequal minima: the primary minimum (Min I) is  $0^{m}.35$  deep, the secondary minimum (Min II) is  $0^{m}.30$  deep. Both minima are total with a duration of  $0^{d}.02$ .

The corresponding ephemeris is:

Min. I = HJD 
$$2451781.392 + 0.2677 \times E$$
.  
  $\pm 0.001 \pm 0.0003$ 

<sup>&</sup>lt;sup>1</sup> Beersel Hills Observatory, Laarheidestraat, 3, 1650 Beersel, Belgium, email: Paul.VanCauteren@advalvas.be

<sup>&</sup>lt;sup>2</sup> Vereniging Voor Sterrenkunde, Belgium

 $<sup>^\</sup>dagger \, {\rm The \ Mira \ AP}$  software is distributed by Axiom Research Inc.

2 IBVS 5057

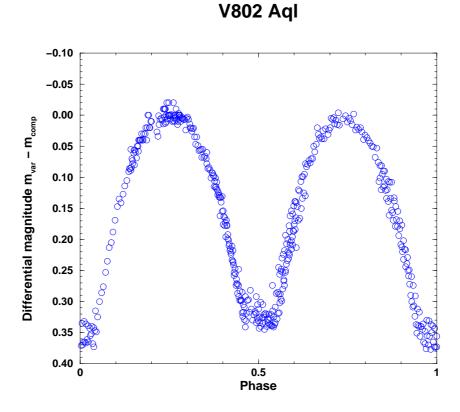


Figure 1. Phase diagram for V802 Aql (against the period of 0d2677)

**Acknowledgement:** This research has made use of the SIMBAD data base operated at the *Centre de Données astronomiques de Strasbourg*, France. We thank Dr. Patricia Lampens, Koninklijke Sterrenwacht van België, for assistance with the text.

## References:

Bakos, G.A., 1950, Ann. Sterrew. Leiden, **20**, No. 4 Rodríguez, E., López-González, M.J., López de Coca, P., 2000, A&AS, **144**, 469 Stellingwerf, R.F., 1978, ApJ, **224**, 953