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**BV PHOTOMETRY OF THE BINARY STAR VW LMi**

DUMITRESCU, A.

Astronomical Institute of the Romanian Academy, Str. Cutitul de Argint 5, RO-75212, Bucharest 28, Romania,  
 e-mail: alex@roastro.astro.ro

<b>Name of the object:</b>	
VW LMi = HIP 54003 = HD 95660	
<b>Equatorial coordinates:</b>	<b>Equinox:</b>
R.A. = 11 <sup>h</sup> 02 <sup>m</sup> 51 <sup>s</sup> .9    DEC. = 30°24'54".5	2000.0
<b>Observatory and telescope:</b>	
Astronomical Institute of the Romanian Academy, 0.5-m <i>f</i> /15 Cassegrain reflector	
<b>Detector:</b>	EMI 9502B type photomultiplier
<b>Filter(s):</b>	<i>B</i> and <i>V</i>
<b>Comparison star(s):</b>	HD 95977, SAO 62390
<b>Check star(s):</b>	HD 94883
<b>Transformed to a standard system:</b>	No
<b>Availability of the data:</b>	
Upon request	
<b>Type of variability:</b>	EW
<b>Remarks:</b>	
<p>VW LMi (HIP 54003, HD 95660) was found to be a variable star by the Hipparcos/Tycho mission. The Variability Annex of the Hipparcos Catalogue (ESA 1997) reports VW LMi to have a period of 0<sup>d</sup>.477547, with Hp magnitude ranging between 8.031 to 8.446. The star was classified as a W UMa eclipsing binary type. The spectral type is listed as F3V. Up to the present, no observations can be found in literature. Nine nights of data were obtained between 15 April 1999 and 9 May 2000. The filters used are in close accordance with the <i>UBV</i> Johnson system. From the light curves produced, four times of minimum light (one primary and three secondaries) were determined. These times, given in Table 1, are calculated according to the method described by Kwee and van Woerden (1956). The phased light curves were obtained using the ephemeris</p> <p style="text-align: center;">Min. I = HJD 2451635.3084 + 0<sup>d</sup>.477547 × <i>E</i>.</p>	

Table 1

Min HJD 2400000.0 +	Type	Epoch
51288.3670	Secondary	-726.5
51617.3998	Secondary	-37.5
51635.3084	Primary	0.
51671.3594	Secondary	75.5

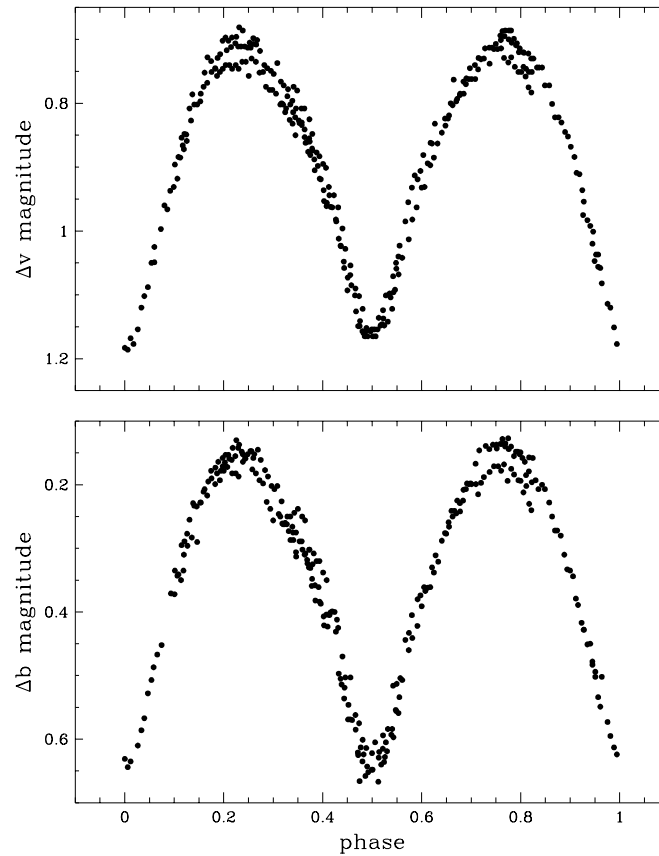


Figure 1.

References:

Kwee, K.K. and van Woerden, H, 1956, *BAN*, **12**, 327  
 ESA, 1997, The Hipparcos Catalogue, ESA SP-1200