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## INFORMATION BULLETIN ON VARIABLE STARS

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## GSC 4832.400: A NEW ECLIPSING BINARY SYSTEM

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| Name of the object: |
| :--- |
| GSC 4832.400 |


| Equatorial coordinates: | Equinox: |
| :--- | :--- |
| R.A. $=07^{\mathrm{h}} 50^{\mathrm{m}} 45.4$ DEC. $=-00^{\circ} 00^{\prime} 100^{\prime \prime} 9$ | 2000.0 |

## Observatory and telescope:

Observatorio del Departamento de Física de la Universidad de Extremadura, Re-
flector Newton $0.4-\mathrm{m} f / 4.5$

| Detector: |
| :--- |
| Filter(s): Starlight Xpress CCD Camera (based in the chip SONY <br> ICX027BL $6.4 \times 4.35 \mathrm{~mm}^{2}, 500 \times 256$ pixels) <br> Comparison star(s): GSC 4832.2073 <br> Check star(s): GSC 4832.912 |


| Transformed to a standard system: | No |
| :--- | :--- |


| Availability of the data: |
| :--- |
| Upon request |

## Type of variability: EW

Table 1

| Min HJD | Type | Epoch | $O-C$ |
| :--- | :--- | :---: | ---: |
| $2451000+$ |  |  |  |
| 1242.3875 | Secondary | 0.5 | 0.0008 |
| 1243.3266 | Secondary | 3.5 | 0.0005 |
| 1244.4210 | Primary | 7 | -0.0012 |
| 1254.4450 | Primary | 39 | 0.0020 |
| 1256.3233 | Primary | 45 | 0.0014 |



Figure 1. The $V$ light curve obtained for GSC 4832.400. Magnitude differences (variable minus comparison) are plotted versus phase, where the phases are computed using the ephemeris calculated in this work.

## Remarks:

The result of this surveillance program showed that GSC 4832.400 is an eclipsing binary system with a period very close to 7.5 hours. Figure 1 shows the differential light curve obtained in the $V$ band. This light curve suggests that GSC 4832.400 could be a near contact binary system. The primary minimum shows $0^{\mathrm{m}} 43$ average depth, and the secondary minimum $0{ }^{\mathrm{m}} 37$. The light curve in the $V$ band also seems to show an O'Connell effect (O'Connell 1951), that amounts to $\Delta m=$ Max. I - Max. $\mathrm{II}=-0.035$ magnitudes, where Max. I is at phase 0.25 and Max. II at phase 0.75.
Five moments of minima (two secondaries and three primaries) were obtained from our observations according to the Kwee-Van Woerden (1956) method. The following ephemeris was derived for the minimum I:

$$
\begin{gathered}
\text { Min. } \mathrm{I}=\mathrm{HJD} 2451242.23011+0 \mathrm{~d} 31315 \times E . \\
\pm 0.00067 \pm 0.00001
\end{gathered}
$$

The times of minima are presented in Table 1. The number of cycles elapsed (E) and $O-C$ residual values are also listed, determined using the ephemeris given above.

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## References:

Kwee, K.K., Van Woerden, H., 1956, BAN, 12, 327

