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**THE NEW SHORT PERIOD EB ECLIPSING BINARY SYSTEM NSV 05339**

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<b>Name of the object:</b>	
NSV 05339, BD +36°2217, CSV 006866, Wr 007, GSC 2526_1034	
<b>Equatorial coordinates:</b>	<b>Equinox:</b>
R.A.= 11 <sup>h</sup> 47 <sup>m</sup> 50 <sup>s</sup> DEC.= +35°13'43"	2000.0
<b>Observatory and telescope:</b>	
Mollerusa Private Observatory, 0.26-m Schmidt–Cassegrain telescope; Esteve Duran Observatory, 0.2-m Schmidt–Cassegrain telescope	
<b>Detector:</b>	CCD
<b>Filter(s):</b>	V
<b>Comparison star(s):</b>	SAO 62705, PPM 76080, GSC 2526_2379
<b>Check star(s):</b>	GSC 2526_781
<b>Transformed to a standard system:</b>	No
<b>Availability of the data:</b>	
Upon request	
<b>Type of variability:</b>	EB
<b>Remarks:</b>	
<p>NSV 05339 was announced to be a variable star by Weber (1955). In the NSV catalogue (Kholopov, 1982) this star is listed as an EA: with a G5 spectral type and a photographic brightness variation between 11<sup>m</sup>0 and 11<sup>m</sup>6. Observations performed between March 1998 and April 1999 show that it is a short period EB eclipsing binary system (Figure 1). About 20' to the east of NSV 05339 is GSC 2506_775, a star which according to GSC has a photovisual magnitude of 13.34. This object was included in the synthetic aperture photometric measurements. Taking this into account, a maximum V amplitude of 0.56 between minimum I and phase 0.25 was obtained. During the secondary minimum the variable fades 0.15 magnitudes, and it also presents an O'Connell effect of 0<sup>m</sup>06. The computed ephemeris is:</p> $\text{Min. I} = \text{HJD } 2451220.4869 + 0^{\text{d}}351862 \times E.$ $\pm 0.0002 \quad \pm 0.000003$	

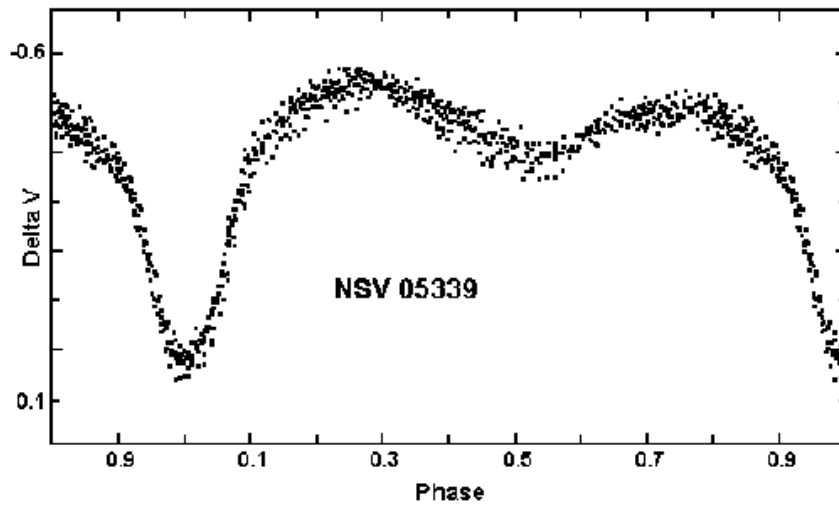


Figure 1.

References:

- Kholopov, P. N., ed., 1982, New General Catalogue of Suspected Variable Stars, Moscow  
Weber, R., 1955, BSAF, 69, 440