# COMMISSIONS 27 AND 42 OF THE IAU INFORMATION BULLETIN ON VARIABLE STARS

Number 4736

Konkoly Observatory Budapest 21 July 1999 *HU ISSN 0374 - 0676* 

## PHOTOMETRY OF THE 1999 OUTBURST OF U SCORPII

#### SEIICHIRO KIYOTA

Variable Star Observers League in Japan, 1-401-810 Azuma, Tsukuba, 305-0031 Japan, e-mail: skiyota@abr.affrc.go.jp

## Name of the object: U Sco

Equatorial coordinates:	Equinox:
<b>R.A.</b> = $16^{h}22^{m}31^{s}2$ <b>DEC.</b> = $-17^{\circ}52'41''$	2000.0

**Observatory and telescope:** 25-cm Schmidt Cassegrain (F/6.3)

Detector:	Apogee AP-7 CCD camera (SITe SIA502AB, $512 \times 512$ )	
		Î

**Filter(s):** Johnson-Cousins  $BVR_c$ 

Comparison star(s):	$GSC 6206_{320} = BD - 17^{\circ}4559 = PPM 23163, V = 10.68,$
	B - V = 0.496 (based on Tycho catalog), $V - R = 0.25$
	(calculated from $B - V$ value)

Transformed to a standard system:

No

Availability of the data: Through IBVS Web-site as 4736-t1.txt

Type of variability: NR

### Remarks:

Patrick Schmeer (1999) reported the 6th recurrent outburst of U Sco (9.5 m<sub>V</sub> on Feb. 25.1940, 1999) (Vsnet-alert 2688, IAUC 7113). L. Shaw reported that U Sco reached the maximum of 7.6 m<sub>V</sub> at Feb. 25.562 (IAUC 7113). I began to observe this outburst immediately after maximum. The lightcurve consists of 3 parts; the 1st fast fading phase, plateau around Mar. 12–22, the 2nd fast fading. B - V was steeply decreasing after maximum but almost constant in the plateau phase. The plateau around 15 days after the maximum was also reported at the outburst during 1987 (Sekiguchi et al., 1988).



Figure 1.

Acknowledgements:	
I thank Dr. Taichi Kato, Kyoto University for useful discussion and advice.	

References:

Schmeer et al., 1999, IAUC 7113
Sekiguchi, K., Feast, M.W., Whitelock, P.A., Overbeek, M.D., Wargau, W., Jones, J.S., 1988, MNRAS, 234, 281