## COMMISSIONS 27 AND 42 OF THE IAU INFORMATION BULLETIN ON VARIABLE STARS

# PHOTOELECTRIC AND CCD TIMES OF MINIMA OF 19 ECLIPSING BINARY SYSTEMS 

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We present photoelectric and CCD photometric minima observations of 19 eclipsing binary systems. Most of them are stars with apsidal motion or at least eccentric binary systems, selected from the listing of Hegedüs (1988). Some minima observations (e.g. for UZ Leonis) are part of complete light curve coverages.

One part of the photoelectric observations was carried out at Piszkéstető Mountain Station of the Konkoly Observatory of the Hungarian Academy of Sciences with a 20 in. $\mathrm{f} / 15$ Cassegrain telescope. The photometer used was equipped with an unrefrigerated EMI9058QB photomultiplier tube and Schott UG2 (for U), BG12+GG13 (for B) and GG11 (for V) filters. This system is referred to as Pi50 in Table 1. The other part of the photoelectric measurements was made at Baja Astronomical Observatory with a Starlight-I photometer, equipped with an unrefrigerated EMI9924A multiplier tube and Schott filters UG1 (matching Johnson's U), GG400+BG25 (matching B) and OG515 (matching V), mounted on the 20 in. f/8.4 Ritchey-Chrétien telescope (Ba50 in Table 1). The unfiltered CCD observations were carried out with an ST-5 (1995-96), and an ST-7 (1997) camera installed on the 20 in . RC telescope mentioned above.

The observations were made between 1995-1998. Reduction of the photoelectric data was made by standard procedures. For the reduction of the CCD frames we used the IRAF package. All the minima times were computed using the parabolic fitting method.

Table 1 presents the derived minima times. The content of the first two columns is self-explaining. The error in the last digit appears in the third column. In the fourth column the types of minima are marked (I for primary, and II for secondary ones), while in the fifth column the number of individual data involved in the parabolic fit is given. The columns from sixth to eighth describe the filters used, the first three letters of the observers' names and the codes of the instrumentation. The last column contains the comparisons used, identified by their BD, GSC or HD numbers.

Table 1

| Star | $\begin{aligned} & \hline \text { Min. HJD } \\ & +2400000 \end{aligned}$ | $\begin{gathered} \text { error } \\ \pm \end{gathered}$ | Min. type | Points used | Filter | Obs.'s name | Instr. | Comp. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| AS Cam | 50519.5178 | 2 | II | 78 | - | Bor | Ba50 (ST7) | GSC 4347-0466 |
| RZ Cas | 50247.5014 | 8 | I | 91 | - | Bor | Ba50 (ST5) | GSC 4317-1437 |
|  | 50815.2545 | 1 | I | 51 | V | Par + Bor | Pi50 | $\mathrm{BD}+67^{\circ} 215$ |
|  | 50815.2541 | 6 | I | 48 | B | $\mathrm{Par}+$ Bor | Pi50 |  |
| PV Cas | 50244.4443 | 3 | II | 134 | - | Bor | Ba50 (ST5) | GSC 4010-1545 |
|  | 50279.4665 | 1 | II | 40 | V | Bir | Pi50 | $\mathrm{BD}+58^{\circ} 2555$ |
|  | 50279.4667 | 4 | II | 40 | B | Bir | Pi50 |  |
| VW Cep | 50171.4028 | 3 | I | 18 | V | Bor | Pi50 | $\mathrm{BD}+75^{\circ} 765$ |
|  | 50171.4030 | 4 | I | 18 | B | Bor | Pi50 |  |
|  | 50171.402 | 1 | I | 19 | U | Bor | Pi50 |  |
|  | 50171.5410 | 3 | II | 27 | V | Bor | Pi50 |  |
|  | 50171.5418 | 1 | II | 27 | B | Bor | Pi50 |  |
|  | 50171.543 | 1 | II | 27 | U | Bor | Pi50 |  |
| XX Cep | 50297.445 | 1 | I | 71 | V | Bir | Pi50 | $\mathrm{BD}+63^{\circ} 2030$ |
|  | 50297.446 | 1 | I | 68 | B | Bir | Pi50 |  |
| GK Cep | 50210.451 | 2 | I | 146 | - | Bir | Ba50 (ST5) | GSC 4465-1764 |
|  | 50225.4286 | 5 | I | 81 | - | Bir | Ba50 (ST5) |  |
| MR Cyg | 50230.4597 | 2 | I | 104 | - | Bor | Ba50 (ST5) | GSC 3609-1293 |
| V453 Cyg | 50235.4809 | 5 | I | 628 | - | Bir | Ba50 (ST5) | GSC 2683-3541 |
| V477 Cyg | 50237.4445 | 2 | I | 120 | - | Bir | Ba50 (ST5) | GSC 2674-0910 |
| V836 Cyg | 49919.4008 | 5 | I | 34 | - | Par | Ba50 (ST5) | GSC 2715-0017 |
| V1136 Cyg | 50270.4672 | 4 | I | 83 | - | Bir | Ba50 (ST5) | GSC 2150-3445 |
| AK Her | 50275.4763 | 3 | I | 47 | V | Bor | Ba50 | $\mathrm{BD}+16^{\circ} 3123$ |
|  | 50275.474 | 1 | I | 48 | B | Bor | Ba50 |  |
|  | 50310.4582 | 5 | I | 32 | V | Bor | Ba50 |  |
|  | 50310.4574 | 2 | I | 32 | B | Bor | Ba50 |  |
|  | 50508.574 | 1 | I | 102 | - | Bor | Ba50 (ST7) | GSC 1536-0928 |
|  | 50512.5802 | 2 | II | 83 | - | Bor | Ba50 (ST7) |  |
| CC Her | 49876.4996 | 3 | I | 38 | - | Par | Ba50 (ST5) | GSC 0946-1166 |
| DI Her | 50238.4879 | 5 | II | 135 | - | Bir | Ba50 (ST5) | GSC 2109-1273 |
| UV Leo | 50499.4473 | 4 | I | 44 | V | Bor | Pi50 | $\mathrm{BD}+14^{\circ} 2277$ |
|  | 50499.4471 | 2 | I | 44 | B | Bor | Pi50 |  |
|  | 50513.5375 | 3 | II | 33 | - | Bor | Ba50 (ST7) | GSC 0845-0136 |
| UZ Leo | 50507.445 | 2 | I | 18 | V | Heg | Pi50 | $\mathrm{BD}+14^{\circ} 2279$ |
|  | 50509.606 | 1 | II | 27 | V | Heg | Pi50 |  |
|  | 50509.607 | 1 | II | 26 | B | Heg | Pi50 |  |
|  | 50510.5361 | 9 | I | 16 | V | Heg | Pi50 |  |
|  | 50510.5386 | 6 | I | 16 | B | Heg | Pi50 |  |
|  | 50512.3872 | 1 | I | 24 | V | Heg | Pi50 |  |
|  | 50512.388 | 1 | I | 21 | B | Heg | Pi 50 |  |
| FT Ori | 50494.3442 | 5 | I | 36 | V | Bor | Pi50 | $\mathrm{BD}+22^{\circ} 1250$ |
|  | 50494.3446 | 3 | I | 36 | B | Bor | Pi50 |  |
| ST Per | 50813.330 | 1 | I | 54 | V | Bor+Par | Pi50 | HD 18615 |
|  | 50813.3301 | 6 | I | 47 | B | Bor+Par | Pi50 |  |
| U Sge | 50287.5178 | 9 | I | 79 | V | Bor | Ba50 | $\mathrm{BD}+19^{\circ} 3976$ |
|  | 50287.5184 | 7 | I | 79 | B | Bor | Ba50 |  |

Remarks on some of the variables:
$X X$ Cep: This Algol-system is an apsidal motion candidate star (e.g. Hegedüs, 1988), but recently Borkovits and Hegedüs (1996) tried to explain its period variation by light-time effect. The new times of minima (see also Hegedüs et al. 1996) do not support their solution.
AK Her: The new times of minima are also inconsistent with the light-time solution of Borkovits and Hegedüs (1996).
CC Her: This minimum time was already published in Hegedüs et al. (1996), but there we missed the correction for Daylight Saving Time (DST), and the GSC number of the comparison was also incorrect. These are the correct data.

PV Cas: The initial comparison candidate star for PV Cassiopeiae, GSC 4010-1432, showed variations of about 0.8 mag against the check star GSC 4010-1545, and some other fainter stars in the CCD frame.
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## References:

Borkovits T., Hegedüs T., 1996, A\&ApSS,120, 63
Hegedüs T., 1988, CDS Bull., 35, 15
Hegedüs T., Bíró I. B., Borkovits T., Paragi Zs., 1996, IBVS No. 4340

## ERRATA

In IBVS Nos. 4555 and 4633 we presented CCD photometric minima observations (together with photoelectric ones) of several eclipsing binary systems. Due to an unfortunate programming bug most of the minimum times have an error in the third decimal place of JD. This erratum contains the corrected moments of minima. Table 1 shows the corrigenda to IBVS No. 4555. Table 2 should be used as a total replacement of the Table of IBVS No. 4633.

Table 1

| Star | Min. HJD <br> +2400000 | Star | Min. HJD <br> +2400000 |
| :--- | :--- | :--- | :--- |
| AS Cam | 50519.5238 | V453 Cyg | 50235.4843 |
| PV Cas | 50244.4435 | V477 Cyg | 50237.4480 |
| GK Cep | 50210.453 | V1136 Cyg | 50270.4694 |
|  | 50225.4297 | DI Her | 50238.4929 |
| MR Cyg | 50230.4608 | UV Leo | 50513.5487 |

Table 2

| Star | Min. HJD <br> +2400000 | error <br> $\pm$ | Min. <br> type | Points <br> used | Filter | Obs.'s <br> name | Instr. | Comp. |
| :--- | :--- | :--- | :---: | :---: | :---: | :---: | :--- | :--- |
| RT And | 50964.5050 | 2 | I | 51 | V | Bir | Ba50 | HD 218915 |
| AB And | 50966.5525 | 3 | II | 30 | V | Bir | Ba50 | GSC 2763-0683 |
|  | 50984.4721 | 4 | II |  | V | Bir | Ba50 |  |
|  | 51016.5005 | 1 | I | 41 | V | Bor | Ba50 |  |
| OO Aql | $50950.486:$ | 1 | I | 19 | V | Bor | Ba50 | HD 187146 |
|  | $50956.5658:$ | 2 | I | 61 | V | Bir | Ba50 |  |
|  | 50967.4659 | 1 | II | 45 | V | Bor | Ba50 |  |
| Y Cam | 50872.4672 | 3 | I | 77 | - | Bor | Ba50 | GSC 4527-1983 |
| AS Cam | 50900.351 | 1 | II | 163 | V | Bor | Ba50 | GSC 4347-0466 |
| RZ Cas | 50871.4318 | 3 | I | 414 | - | Bor | Ba50 | GSC 4317-1578 |
| TV Cas | $51005.45:$ | 1 | II | 190 | V | Bir | Ba50 | GSC 3665-0026 |
| PV Cas | 51015.5244 | 5 | I | 55 | V | Bir | Ba50 | GSC 4010-1432 |
| VW Cep | 50871.6279 | 5 | I | 102 | - | Bor | Ba50 | GSC 4585-2387 |
|  | 50900.5736 | 3 | I | 54 | V | Bor | Ba50 |  |
|  | 50941.3443 | 3 | II | 51 | V,B,R | Bor | Ba50 |  |
|  | 50941.4859 | 3 | I | 61 | V,B,R | Bor | Ba50 |  |
|  | 50942.4573 | 5 | II | 59 | V,B,R | Bir | Ba50 |  |
|  | 50942.5990 | 5 | I | 28 | V,B,R | Bir | Ba50 |  |
| XX Cep | $51018.517:$ | 6 | II | 91 | V | Bor | Ba50 | GSC 4288-0186 |
| CQ Cep | 50948.5431 | 6 | I | 200 | V | Bir | Ba50 | GSC 3991-1316 |
| DL Cyg | 51038.485 | 3 | I | 170 | V,R | Bor | Ba50 | GSC 3595-0816 |
| MR Cyg | 50962.4954 | 1 | II | 87 | V | Bir | Ba50 | GSC 3609-1220 |
|  | 51014.4830 | 5 | II | 124 | V | Bir | Ba50 |  |
| V477 Cyg | 50974.4054 | 2 | I | 53 | V | Bor | Ba50 | GSC 2674-0910 |
| AK Her | 50865.6038 | 2 | I | 53 | V,B | Bir | Pi50 | BD+16³123 |
|  | 50866.6601 | 1 | II | 78 | - | Bor | Ba50 | GSC 1536-0928 |
|  | 50884.5722 | 2 | I | 82 | V | Bor | Ba50 |  |
|  | 50903.5413 | 3 | I | 78 | V | Bir | Ba50 |  |
|  | 50971.4060 | 3 | I | 266 | R | Bor | Ba50 |  |

Table 2 (cont.)
$\left.\begin{array}{llcccccll}\hline \text { Star } & \begin{array}{l}\text { Min. HJD } \\ +2400000\end{array} & \begin{array}{c}\text { error } \\ \pm\end{array} & \begin{array}{c}\text { Min. } \\ \text { type }\end{array} & \begin{array}{c}\text { Points } \\ \text { used }\end{array} & \text { Filter } & \text { Obs.'s } & \text { Instr. Comp. } \\ \text { name }\end{array}\right]$
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