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TWO NEW VARIABLE STARS

My continuing photographic patrol has resulted in the discovery of 2 more variable stars that are not listed in the General Catalogue of Variable Stars (Kholopov et al., 1985) or the subsequent Name Lists of Variable Stars (Kholopov et al., 1985, 1987, 1989; Kazarovets and Samus, 1990, Kazarovets et al., 1993; Kazarovets and Samus, 1995). Nor are they listed in the New Catalogue of Suspected Variable Stars (Kholopov et al., 1982). Positions and preliminary magnitude ranges, types, and periods are given in Table 1, which continues the list in Kaiser (1994). Figures 1 and 2 are finder charts for DHK 42 and 43 respectively.

Table 1

Var. Designation	RA (1950)	Dec (1950)	Range	Туре	P (days)
DHK 42=GSC 1272:0567	$4^{\rm h}18^{\rm m}29^{\rm s}$	20°8′55″	12.5-<15.5p	Μ	\sim 322
DHK 43=GSC 2759:1984	$23^{\rm h} \ 8^{\rm m} \ 1^{\rm s}$	34°30′1″	11.7 - 12.4p	SR	~ 60

Both variables were discovered using photographs taken with the Automatic Photographic Patrol (APP). David B. Williams has examined both DHK 42 and 43 on the Harvard patrol plates and has confirmed variability and determined preliminary periods. DHK 42 was also visually confirmed by Tim Hager, who reported it as <14.2 in March of 1995 and as bright as 11.3 in August, 1995. It is a mira variable located in Taurus on the far side of the Hyades star cluster. DHK 43 is a semiregular variable in Pegasus. Both photographic and ongoing photoelectric observations will be published when analysis is completed.

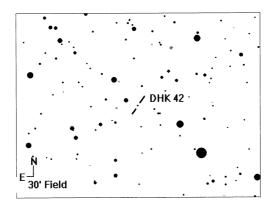


Figure 1

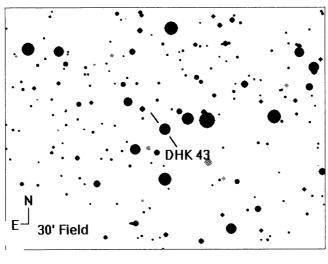


Figure 2

The construction of the APP was supported by a grant from NASA administered by the American Astronomical Society to whom I'm most grateful. I would also like to thank David B. Williams and Tim Hager for helping confirm the variability of these stars. The variables were detected using a projection blink comparator (PROBLICOM) as described by Mayer (1977).

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