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## X PERSEI

The Be/X-ray binary X Persei is the optical counterpart of X-ray pulsar 4U $0352+30$. Possible orbital period of the system is the 580 day periodicity detected in the radial velocities of the Balmer lines (Hutchings, 1977), in spite of existence of some doubts (Penrod \& Vogt, 1985). The star is a known optical variable on time scales from minutes to years. The most comprehensive optical light curve over the past decades is presented by Roche et al. (1993).

X Per has been observed with 60 cm telescope and photon counting photometer of the National Astronomical Observatory Rozhen. HR 1197 ( $V=6.25, \mathrm{~B}-\mathrm{V}=0.20, \mathrm{U}-\mathrm{B}=0.14$ ) served as the comparison star. The APR software (Kirov et al., 1993) was used during the data reduction.

Our data are summarized in Table 1. Figure 1 presents the $V$ band light curve over the period 1986-1995. Our observations showed the optical low state that began in the mid-1990 finished in the spring of 1993. Since this time the star has been in optical high state. A clear similarity between the light maximum in 1987-1988 and that in 1993-1994 is visible. The displacement is about 2400 days, i.e. 4 times larger than the supposed orbital period.

The last observations indicate that the optical high state is almost over and X Per is going to a new low state now. May be the star will lose its circumstellar disk again as it was observed in 1990 (Norton et al., 1991; Roche et al., 1993).


Figure 1. V band light curve of X Per. Circles indicate our data. The light curve before JD 2448500 is taken from Roche et al. (1993).

Table 1

| JD 2440000+ | V | $\mathrm{B}-\mathrm{V}$ | $\mathrm{U}-\mathrm{B}$ | $\mathrm{JD} 2440000+$ | V | $\mathrm{B}-\mathrm{V}$ | $\mathrm{U}-\mathrm{B}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  |
| 8545.57 | 6.81 | 0.09 | -0.68 | 9341.27 | 6.47 | 0.18 | -0.78 |
| 8620.52 | 6.76 | 0.08 | -0.64 | 9343.42 | 6.47 | 0.19 | -0.78 |
| 8712.29 | 6.78 | 0.11 | -0.73 | 9357.43 | 6.44 | 0.20 | -0.78 |
| 8734.26 | 6.76 | 0.05 | -0.68 | 9366.46 | 6.44 | 0.19 | -0.78 |
| 8841.52 | 6.75 | 0.13 | -0.66 | 9572.51 | 6.34 | 0.26 | -0.73 |
| 8842.56 | 6.76 | 0.11 | -0.67 | 9573.52 | 6.33 | 0.26 | -0.73 |
| 8905.56 | 6.78 | 0.09 | -0.70 | 9597.50 | 6.26 | 0.26 | -0.73 |
| 8916.42 | 6.78 | 0.09 | -0.70 | 9598.48 | 6.27 | 0.24 | -0.73 |
| 9028.22 | 6.78 | 0.05 | -0.66 | 9715.32 | 6.38 | 0.21 | -0.75 |
| 9046.26 | 6.76 | 0.06 | -0.66 | 9748.23 | 6.45 | 0.21 | -0.81 |
| 9195.55 | 6.45 | 0.21 | -0.75 | 9772.23 | 6.48 | 0.21 | -0.83 |
| 9197.54 | 6.42 | 0.21 | -0.76 |  |  |  |  |
| 9340.28 | 6.47 | 0.18 | -0.79 |  |  |  |  |

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