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## POSITIONS OF VARIABLE STARS IN PLAUT'S FIELD 3

This paper continues the study announced in our previous publication (Antipin et al., 1994a) and deals with the variable stars in the Palomar-Groningen Field 3 (Plaut, 1971). This field contains 1474 variable stars discovered by Plaut. Using photographic finding charts prepared by L. Plaut and available in Moscow, we have checked for possible GSC identifications. 174 stars have been identified with the GSC. We have also checked 83 stars discovered in this field by other authors and identified 23 of them with GSC objects. We are not able to present the corresponding table here because of volume restrictions and are planning to publish this list in Astronomical and Astrophysical Transactions. We have not found very significant positional mistakes among those stars we were able to identify with GSC. We have also checked positions for non-GSC variable stars and found quite a number of serious mistakes in GCVS positions. The corrected coordinates presented in

Table 1. Corrected Coordinates for Variable Stars

| Plaut | GCVS, NSV | $\alpha(2000.0)$ | $\delta(2000.0)$ | $\Delta \alpha$ | $\Delta \delta$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 23 | V2636 Sgr | $18^{\mathrm{h}} 12^{\mathrm{m}} 23.0$ | $-31^{\circ} 20^{\prime} 40^{\prime \prime}$ | $-0 .{ }^{\mathrm{m}} 01^{\text {s }}$ | $-1^{\circ} 00.0$ |
| 34 | V2647 Sgr | 181238.1 | -310800 | -0 03 | -0 00.6 |
| 111 | V2705 Sgr | 181349.7 | -34 2046 | -0 02 | -0 00.8 |
| 307 | V2852 Sgr | 181643.2 | -34 0122 | +0 00 | -0 03.4 |
| 437 | V2965 Sgr | 181827.0 | -32 1039 | -0 01 | -0 11.2 |
| 474 | V2988 Sgr | 181142.2 | -36 1144 | +0 10 | -0 01.5 |
| 521 | V3033 Sgr | 181232.2 | -32 3915 | -0 42 | -0 04.0 |
| 580 | V3081 Sgr | 182036.2 | -34 4621 | +0 07 | +0 00.2 |
| 707 | V3180 Sgr | 182302.2 | -33 5725 | +0 00 | -0 00.9 |
| 732 | V3200 Sgr | 182326.6 | -30 5407 | -0 02 | -0 17.6 |
| 755 | V3220 Sgr | 182349.5 | -31 1143 | -0 02 | -0 16.1 |
| 782 | V3243 Sgr | 182446.0 | -32 0301 | -0 32 | +0 00.1 |
| 835 | V3288 Sgr | 182515.0 | -33 3346 | +001 | +0 11.3 |
| 881 | NSV10777 | 182613.4 | -314159 | +0 04 | +0 00.1 |
| 1027 | V3447 Sgr | 182113.8 | -30 5346 | +842 | -0 00.8 |
| 1030 | V3443 Sgr | 182956.2 | -30 5737 | +0 01 | -3 44.3 |
| 1158 | V3549 Sgr | 183313.3 | -31 2227 | +0 00 | -0 11.2 |
| 1388 | V3723 Sgr | 183939.1 | -34 4116 | +0 14 | -152.0 |
| 1395 | V3729 Sgr | 183816.7 | -365146 | +152 | +0 00.6 |
| 1427 | V3758 Sgr | 183214.2 | -31 0934 | $-100$ | +0 00.1 |
| 1453 | V3774 Sgr | 184141.2 | -32 5435 | +0 04 | -0 33.4 |
| 1474 | V666 CrA | 182222.7 | $-365957$ | -0 09 | +0 00.1 |
|  | NSV 10351 | 181220.9 | -315024 | +015 | +0 01 |
|  | NSV 10364 | 181259.4 | -314647 | +0 13 | +0 00.2 |

Table 2. Identifications of Several GCVS Stars

| 474 | $\mathrm{~V} 2988 \mathrm{Sgr}=\mathrm{V} 655 \mathrm{Sgr}$ | Plaut, 1971 |
| :--- | :--- | :--- |
| 595 | $\mathrm{NSV} 10648=\mathrm{V} 1289 \mathrm{Sgr}$ | Mayall, 1951 (chart) |
| 723 | $\mathrm{~V} 1188 \mathrm{Sgr}=\mathrm{NSV} 10715$ | Strohmeier et al., 1964 |
| 1063 | $\mathrm{~V} 3475 \mathrm{Sgr}=\mathrm{V} 2355 \mathrm{Sgr}$ | Rosino, 1962 |
| 1105 | $\mathrm{~V} 3507 \mathrm{Sgr}=\mathrm{V} 2360 \mathrm{Sgr}$ | Rosino, 1962 |
| 1388 | $\mathrm{~V} 3723 \mathrm{Sgr}=\mathrm{V} 1635 \mathrm{Sgr}$ | Plaut, 1971 |
| 1395 | $\mathrm{~V} 3729 \mathrm{Sgr}=\mathrm{V} 2371 \mathrm{Sgr}$ | Plaut, 1971 |
| 1414 | $\mathrm{~V} 3746 \mathrm{Sgr}=\mathrm{V} 948 \mathrm{Sgr}=\mathrm{V} 1154 \mathrm{Sgr}$ | GCVS |

Table 1 have been determined by one of us (S.A.) relative to GSC stars. The columns of the table contain: Plaut's numbers (the two last lines, without Plaut's numbers, refer to variables discovered by Rosino (1962), in the field of the globular cluster NGC 6569); GCVS or NSV designations; right ascensions and declinations (2000.0); positional differences found (in the sense GCVS minus present study).

Table 2 contains 8 stars for which positional inaccuracies have led to multiple entries in the GCVS and the NSV catalogue. The first column gives Plaut's numbers. The second column presents the existing GCVS or NSV designations. The last column of the table identifies the source of error. If the star in Table 2 does not enter also Table 1, Plaut's coordinates for this star are to be preferred.

Several final remarks deal with difficult cases. It turns out that Plaut's variable No. 239 $\left(18^{\mathrm{h}} 12^{\mathrm{m}} 32^{\mathrm{s}},-31^{\circ} 07!3,1950.0\right.$, Cepheid with a period of 51 : days according to Plaut, 1971) never got a GCVS designation. The reasons for this are presently unclear, no indication of any mistakes revealed is present in GCVS team records. Though the star could not be identified with any GSC star, Plaut's coordinates agree with the photograph. We have not been able to find Plaut's No. $1282=$ V3643 Sgr. The photograph does not agree with the coordinates, a great mistake in coordinates is probable.

Mr. J. Mánek has turned our attention to a mistake in the second paper of the present series (Antipin et al., 1994b). The declination presented there for V448 Oph $=$ GSC 6237.1702 is wrong, the correct $\delta(2000.0)$ value is $-18^{\circ} 06^{\prime} 56^{\prime \prime}$.

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