## COMMISSIONS 27 AND 42 OF THE IAU INFORMATION BULLETIN ON VARIABLE STARS

Number 4128

Konkoly Observatory Budapest 23 December 1994 HU ISSN 0374 - 0676

## B AND V PHOTOELECTRIC OBSERVATIONS OF THE CONTACT BINARY XY Leo

XY Leo is a short period contact binary with multiple CaII H and K emission lines and strong X-ray emission (Hrivnak et al., 1984; Vilhu and Rucinski, 1983). The O-C diagram has a sine-like shape, but sudden changes of period of XY Leo were observed three times (Gehlich et al., 1972). These phenomena were explained successfully by BY Dra type activity of the other pair of stars in the system (Barden, 1987). The variations in the light curves of the binary were detected as usually for all W-systems.

Our observations of XY Leo were carried out on three nights: 28, 29 and 30 November 1990). The four channel WBVR photometer attached to the 48-cm reflector was used (Kornilov and Krylov, 1990). The observations were made at High Altitude Alma-Ata Observatory of Sternberg Astronomical Institute. BD +28°2036 served as the comparison star. The mean error of magnitudes range from  $0^{\text{m}}$ 01 to  $0^{\text{m}}$ 03 in each filter. These values are due to small ratio of signal to noise ( $\sim 3$ ) for stars of such brightness. V and B light curves of XY Leo are presented in Figure 1. Phases are calculated by the ephemeris from Krzesinski et al. (1990):

## $MinI=J.D.hel 2447612.34748+0.2841034\times E$

The observations allowed us to determine the times of one primary and two secondary minima by graphical method separately for each colour. Table I contains the moments of minima, type of minima and O-C values as calculated from the same ephemeris.

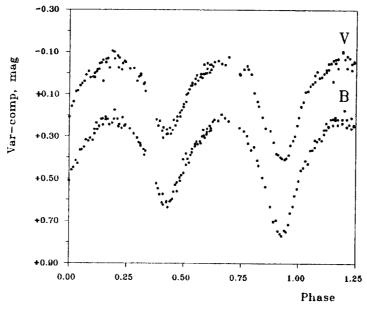


Figure 1

Table I

${ m JD}_{hel} \ 2448000+$	Min	Filter	О-С
224.4477: 224.4470: 225.4393 225.4399 226.4348 226.4351	II II I II	V B V B	$-0.0194 \\ -0.0201 \\ -0.0222 \\ -0.0216 \\ -0.0210 \\ -0.0207$

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