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## PHOTOELECTRIC MINIMA OF ECLIPSING BINARIES

The following Table gives the photoelecric minima obtained in the years 1991-1993 at the N Copernicus Observatory and Planetarium in Brno (Czech Republic) by means of the Nasmyth type $0.4-\mathrm{m}$ telescope.

The telescope was used with a single channel photometer and EMI 6256B photomultiplier. Measurements were made in the UBV-system. The integration time of one measurement was ten second.

The times of minimum brightness and their standard deviation were determined by Kwee-van Woerden's (1956) method. The name of star, the name of the filter, heliocentric time of minima, standard deviation, different values of $\mathrm{O}-\mathrm{C}$ and abbreviations of the observers's name are given in Table 1. The abbreviations in the column "Observer" means:
DH: Dalibor Hanžl EN: Eva Neureiterová PH: Petr Hájek TH: Tomáš Hudeček MN: Martin Navrátil MZ: Miloslav Zejda
Data for calculating the $\mathrm{O}-\mathrm{C}$ residuals have been taken from the following literature: O-C(I): SAC 65, Krakow, 1993
O-C(I): GCVS, Moscow, 1985-1987
Table 1

| Name |  |  | JD hel | err. | O-C(I) |  | O-C(II) | Obs. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| RT And | V | s | 48506.3373 | 0.0016 | -0.0035 |  | -0.0017 | MN |
|  | V |  | 48600.3639 | 0.0005 | -0.0019 |  | -0.0001 | MN |
|  | B |  | 48600.3629 | 0.0003 | -0.0029 |  | -0.0011 | MN |
|  | U |  | 48600.3655 | 0.0004 | -0.0003 |  | -0.0015 | MN |
|  | V |  | 48646.2751 | 0.0014 | -0.0025 |  | -0.0007 | MN |
|  | B |  | 48646.2746 | 0.0008 | -0.0030 |  | -0.0012 | MN |
|  | U |  | 48646.2754 | 0.0004 | -0.0022 |  | -0.0004 | MN |
| DS And |  | V | 48537.4321 | 0.0005 | +0.0050 | $=$ | $+0.0050$ | DH |
|  | B |  | 48537.4293 | 0.0007 | +0.0023 | $=$ | +0.0023 | DH |
| RX Ari | V |  | 48262.3965 | 0.0002 | -0.0073 |  | $+0.0097$ | DH |
|  | B |  | 48262.3951 | 0.0005 | -0.0087 |  | $+0.0083$ | DH |
| TT Aur | V |  | 48599.3018 | 0.0004 | $+0.0073$ |  | -0.0059 | DH |
|  | B |  | 48599.3007 | 0.0002 | $+0.0062$ |  | -0.0070 | DH |
|  | U |  | 48599.3016 | 0.0003 | $+0.0071$ |  | $-0.0061$ | DH |
| BF Aur | V | s | 48271.3680 | 0.0004 | +0.0052 | $=$ | +0.0052 | DH/PH |
|  | B | s | 48271.3694 | 0.0003 | +0.0066 | $=$ | $+0.0066$ | DH/PH |
|  | U | s | 48271.3724 | 0.0011 | +0.0096 | = | $+0.0096$ | DH/PH |
| VW Cep | V |  | 48276.4159 | 0.0005 | -0.0000 |  | -0.0533 | EN |
|  | B |  | 48276.4144 | 0.0005 | -0.0015 |  | -0.0548 | EN |
| GS Cep | V | s | 48461.4968 | 0.0012 | +0.0006 |  |  | DH |
|  | B | s | 48461.4969 | 0.0020 | $+0.0007$ |  |  | DH |
|  | V |  | 48500.4957 | 0.0008 | +0.0015 |  |  | DH |
|  | B |  | 48500.4949 | 0.0008 | $+0.0007$ |  |  | DH |
|  | V |  | 48503.4394 | 0.0011 | +0.0019 |  |  | DH |
|  | B |  | 48503.4397 | 0.0006 | $+0.0022$ |  |  | DH |

Table 1 (cont.)

| Name |  |  | JD hel | err. | O-C(I) | O-C(II) |
| :--- | :--- | :---: | :---: | :---: | :---: | :--- | Obs.

Moments of the secondary minima are labelled by "s". As far as the data for calculating the time of the secondary minima are not given in the above mentioned literature, we use the phase 0.5 for calculating the $\mathrm{O}-\mathrm{C}$ of the secondary minima (the secondary minimum is supposed to be in mid-phase between the primary ones).

In case the elements in both sources are equal, the $\mathrm{O}-\mathrm{C}$ 's are also equal (this is indicated by the sign "="). The GCSV period of GS Cep is definitely wrong, so corresponding $(\mathrm{O}-\mathrm{C})$ values are not given.

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## References:

GCVS, Moscow, 1985-87
Kwee, K. K. and van Woerden, H., 1956, Bull. Astron. Inst. Neth., 12, No. 464
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