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**BVR OBSERVATIONS OF THE DOUBLE-MODE  
CEPHEID BQ Ser**

Recently one of us (Berdnikov, 1992) has decomposed available photoelectric observations of all double-mode Cepheids into two oscillations. It was pointed out as well that the number of the existing observations is not sufficient for certain investigation of light curve variations in the majority of these stars. Therefore we continue to carry out the observations of the all accessible double-mode Cepheids.

Table 1

JD hel 2440000+	V	B-V	V-R	JD hel 2440000+	V	B-V	V-R
9197.3755	9.632	1.547	1.338	9220.2735	9.341	1.375	1.231
9198.2273	9.778	1.590	1.365	9221.3286	9.378	1.440	1.260
9198.3235	9.773	1.566	1.378	9222.2763	9.648	1.572	1.334
9199.3321	9.207	1.318	1.203	9224.2880	9.479	1.472	1.271
9200.2375	9.429	1.442	1.280	9225.2501	9.455	1.430	1.277
9201.2437	9.726	1.571	1.364	9226.2522	9.420	1.425	1.275
9201.3630	9.735	1.586	1.358	9227.2446	9.573	1.539	1.323
9202.2491	9.633	1.542	1.303	9228.2541	9.755	1.583	1.356
9202.3635	9.616	1.512	1.320	9229.2445	9.221	1.330	1.199
9203.2318	9.414	1.414	1.262	9230.2469	9.421	1.458	1.289
9203.3535	9.424	1.433	1.270	9231.2479	9.726	1.595	1.365
9204.2292	9.477	1.456	1.284	9232.2356	9.673	1.533	1.350
9204.3528	9.475	1.459	1.293	9233.2373	9.359	1.389	1.244
9205.2326	9.502	1.462	1.311	9234.2370	9.499	1.486	1.293
9205.3627	9.516	1.474	1.302	9235.2304	9.536	1.507	1.304
9206.2591	9.615	1.539	1.338	9236.2145	9.601	1.505	1.328
9206.3721	9.633	1.551	1.321	9237.2066	9.640	1.524	1.319
9207.2389	9.681	1.515	1.341	9238.1978	9.242	1.356	1.216
9207.3558	9.624	1.486	1.306	9239.2023	9.474	1.485	1.297
9209.2223	9.484	1.489	1.294	9240.2145	9.772	1.609	1.372
9209.3636	9.521	1.505	1.330	9241.2169	9.578	1.489	1.308
9210.2941	9.772	1.606	1.372	9243.1952	9.577	1.529	1.324
9211.2693	9.559	1.476	1.288	9244.2201	9.640	1.545	1.339
9212.2726	9.348	1.396	1.267	9245.2177	9.559	1.496	1.317
9218.2878	9.558	1.503	1.332	9298.1059	9.353	1.382	1.247
9219.2781	9.805	1.589	1.389				

Photoelectric observations of double-mode Cepheid BQ Ser were carried out in summer-autumn 1993. The 60-cm reflector of the Mt. Maidanak observatory of the Tashkent Astronomical Institute was used and 51 BVR measurements (Table 1) were obtained.

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