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**BVR OBSERVATIONS OF THE DOUBLE-MODE
 CEPHEID V367 Sct**

Recently one of us (Berdnikov, 1992) has decomposed available photoelectric observations of all double-mode Cepheids into two oscillations. It was pointed out as well that the number of the existing observations is not sufficient for certain investigation of light curve variations in the majority of these stars. Therefore we continue to carry out the observations of the all accessible double-mode Cepheids.

Table 1

JD hel 2440000+	V	B-V	V-R	JD hel 2440000+	V	B-V	V-R
9197.3653	11.770	1.867	1.755	9220.2672	11.768	1.917	1.757
9198.2323	11.525	1.785	1.677	9221.3206	11.705	1.890	1.711
9198.3152	11.501	1.761	1.664	9222.2706	11.685	1.869	1.733
9199.3218	11.347	1.766	1.616	9224.2812	11.496	1.810	1.677
9200.2285	11.597	1.890	1.734	9225.2417	11.478	1.781	1.678
9201.3534	11.887	-	1.824	9226.2434	11.625	1.852	1.730
9202.2407	11.932	1.926	1.841	9227.2354	11.867	1.918	1.842
9202.3558	11.927	1.994	1.812	9228.2448	11.895	1.987	1.803
9203.2251	11.714	1.854	1.726	9229.2357	11.559	1.825	1.662
9203.3454	11.677	1.905	1.708	9230.2385	11.314	1.707	1.610
9204.2220	11.431	1.765	1.656	9231.2409	11.584	1.813	1.708
9204.3444	11.417	1.757	1.616	9232.2282	11.807	1.932	1.794
9205.2256	11.522	1.832	1.676	9233.2299	11.859	1.927	1.792
9205.3538	11.567	1.811	1.713	9234.2295	11.754	1.874	1.775
9206.2525	11.679	1.885	1.750	9235.2232	11.564	1.843	1.708
9206.3647	11.645	1.932	1.732	9236.2080	11.558	1.795	1.690
9207.2329	11.710	1.859	1.746	9237.1994	11.610	1.808	1.742
9207.3490	11.666	1.923	1.714	9238.1924	11.585	1.796	1.720
9209.2162	11.759	1.896	1.759	9239.1965	11.640	1.852	1.740
9209.3562	11.751	1.965	1.746	9240.2080	11.829	1.909	1.803
9210.2858	11.701	1.823	1.732	9241.2094	11.833	1.908	1.773
9211.2639	11.501	1.770	1.669	9243.1895	11.321	1.734	1.613
9212.2673	11.376	1.761	1.638	9244.2140	11.617	1.856	1.739
9218.2807	11.584	1.837	1.721	9245.2053	11.811	1.972	1.794
9219.2718	11.737	1.891	1.759	9298.0968	11.777	1.947	1.750

Photoelectric observations of double-mode Cepheid V367 Sct were carried out in summer-autumn 1993. The 60-cm reflector of the Mt. Maidanak observatory of the Tashkent Astronomical Institute was used and 50 BVR measurements (Table 1) were obtained.

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