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**H α -PROFILE IN THE SPECTRUM OF PLEIONE IN THE
 BEGINNING OF THE NEW Be PHASE**

Guo Yulian (1993) reported about significant variations of the H α -profile of Pleione (BU Tau) in the beginning of a new Be-phase (December 1991). He draw the conclusion that the H α -profile with double emission peak turned into a single emission peak after December 1991.

Observations of the H α -emission line were also carried out by the authors during the same period (September 1991-January 1992) using the Coudé-spectrograph with CCD of the 2.6 m-telescope of Crimean Astrophysical Observatory. (A detailed description of this CCD-camera is given in Beriozin et al., 1991.) The linear dispersion of the obtained spectra was 6 Å/mm (0.105 Å/pixel), the spectral interval 60 Å. For one spectrum (1991 December 13) the linear dispersion was 3 Å/mm (0.05 Å/pixel), the spectral interval 30 Å. The S/N-ratio at the continuum was >100. The data reduction was done by using the software of Crimean Astrophysical Observatory. Figure 1 illustrates the obtained H α -profiles. All profiles show double emission peaks certainly. In Table 1 the main parameters of the H α -profiles are summarized.

Table 1

No.	Date	W(Å)	FWHM(Å)	I _b	I _r	I _b /I _r	I _{abs}	$\Delta\lambda$ (Å)
1	1991 September 4	28.98	6.92	4.45	4.96	0.90	3.94	2.76
2	1991 December 13	32.86	6.99	5.27	5.38	0.98	4.02	2.81
3	1991 December 14	29.20	6.94	4.70	4.92	0.96	3.91	2.84
4	1992 January - 11	30.94	6.81	4.75	5.38	0.88	4.07	2.77

W -equivalent width relative to the adjacent continuum
 FWHM -full-width-half-maximum intensity of the common H α -profile
 I_b, I_r, I_{abs} -intensities of the blue, red and central absorption components
 relatively to the continuum, respectively
 I_b/I_r -ratio of the intensities of the blue and red peaks
 $\Delta\lambda$ -separation of the blue and read peaks

Our observational data led to the following conclusions:

1. In the beginning of the new Be-phase the double peak of the H α -emission line remained. The single emission peak obtained by Guo Yulian (1993) is obviously the result of the insufficient spectral resolution of his observations (D=1.15Å/pixel).
2. The equivalent width of the H α -emission line does not differ very much from those published by Guo Yulian (W=26Å). The intensity of the H α -emission line of our observations is higher than that of Guo Yulian (I_{max}=3.8).
3. Our observations also confirm that the intensity of the H α -emission line has increased since the beginning of the new Be-phase.

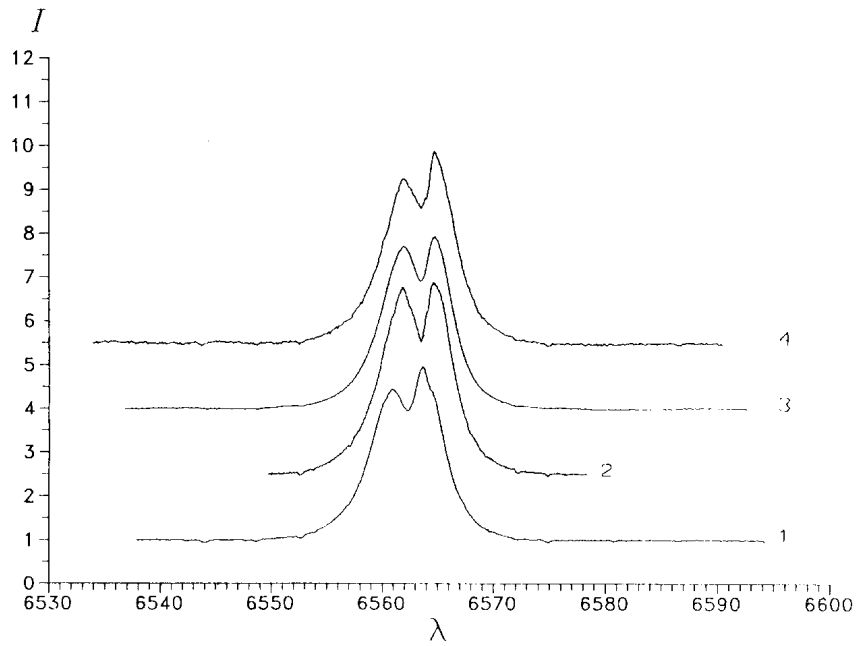


Figure 1. H α -profiles of Pleione.

E. V. MENCHENKOVA
 Astronomical Observatory of the
 Odessa State University
 Park Shevchenko
 270014 Odessa
 Ukraine

R. LUTHARDT
 Sonneberg Observatory
 96515 Sonneberg
 Germany

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