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**NSV 07453 IS A NEW SHORT PERIOD CEPHEID
 OR A LONG PERIOD RR LYRAE STAR**

The variability of NSV 07453=CSV 7263=WR 88 was discovered by Weber (1959). The author pointed out the brightness range 12^m2-13^m6 and the possible type: Cepheid with the period of about 12^d .

The star was studied on 247 photographic plates obtained during J. D. 2418528-48394. The chart and the magnitudes of the comparison stars are given in Fig. 1 and Table 1 respectively. The analysis of the observations revealed that the brightness range is $12^m5 - 13^m8$ pg.

The following elements have been obtained:

$$J. D. \max = 2441828.36 + 0^d929172 \times E, \quad M-m = 0^m2.$$

The average light curve constructed with the new elements is given Fig.2. This period is unusual for such an amplitude. Note that the one-day alias of this period is equal to 13^d3 . This value is very close to the period determined by Weber.

We compiled a sample list from GCSV (Kholopov, 1985) of the stars with the period $0^d87 < P < 1^d0$ and amplitudes $\geq 1^m$. The list of these stars is given in Table 2. These objects are sometimes called the BL Her stars, but BL Her itself has a slightly longer period ($P=1^d307$). There exists a very similar object, FY Vir (Goranskij and Shugarov, 1979). FY Vir has $P=1^d082$ and the range from 15^m7 to 17^m3 pg.

Additional high quality spectroscopic observations are desirable to classify this object correctly.

Table 1

Star	B_{pg}
a	12^m58
b	13.07
c	13.59
d	14.15

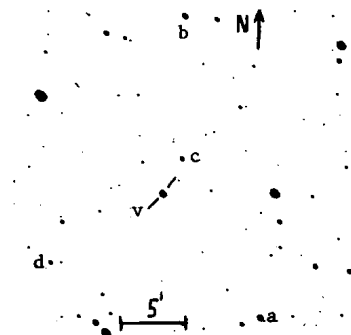


Fig.1

Table 2. Selected GCSV Stars (see text)

Name	Max-Min	Period	Type	M-m
V900 Aql	14 ^m 6 - 16 ^m	0 ^d 874	RR	20:
V524 Her	14.6 - 16.2	0.933	RRAB	11
V742 Cyg	15.9 - 16.7	0.93946	CWB	25
IZ Aps	13.7 - 14.7	0.9775	RRAB	20
EN Pav	15.2 - 16.4	0.9775	RRAB	20

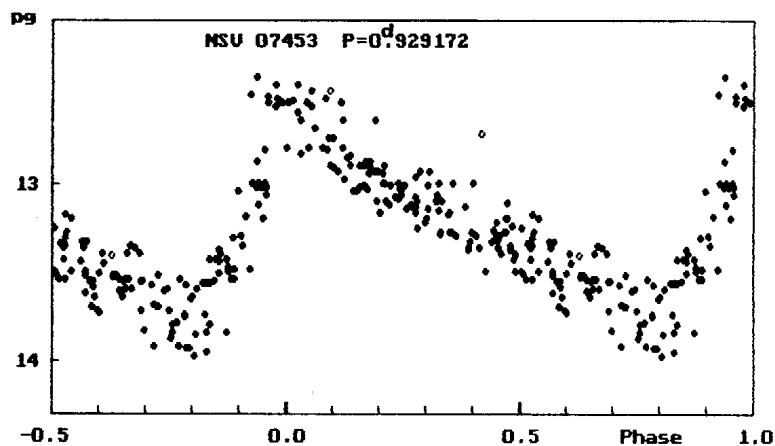


Fig. 2. The mean light curve of NSV 07453

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