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NEW PHOTOELECTRIC MINIMA TIMES OF TZ DRACONIS AND ITS PERIOD STUDY

The eclipsing binary star TZ Dra (BD +47° 2625 (Faulkner, 1986)) has been observed photoelectrically from the National Observatory of Athens, Greece, during 1983. The observations were made using a two-beam, multi-mode, nebular-stellar photometer attached to the 48-inches Cassegrain reflector at the Kryonerion Astronomical Station.

The stars HD 170074 (of $m_v=9.0$) and HD 170357 (of $m_v=8.31$) were used for comparison and checking, respectively. The two intermediate pass bands of the filters used are in close accordance to the international UBV system and the photometer was cooled using dry ice. Reduction of the observations has been made (Hardie, 1962) as usual.

From our observations 4 new minima times (three primaries and one secondary) were derived. They were found using Kwee and Van Woerden's method (1956) and are the mean values of our B and V observations. They are presented in Table I the successive columns of which give: the Heli. JD; the residuals $(O-C)_1$, $(O-C)_{11}$ and $(O-C)_{111}$; and the corresponding number of cycles passed, E_1 , E_{11} and E_{111} . The C's have been calculated using the three different proposed ephemeris formulae for TZ Draconis, which are:

$$\text{Min I} = 2433871.389 + 0^d.8660337 E \quad (\text{I})$$

(due to Perova, 1952)

$$\text{Min I} = 2437911.4347 + 0^d.8660333 E \quad (\text{II})$$

(due to Plavec, 1964)

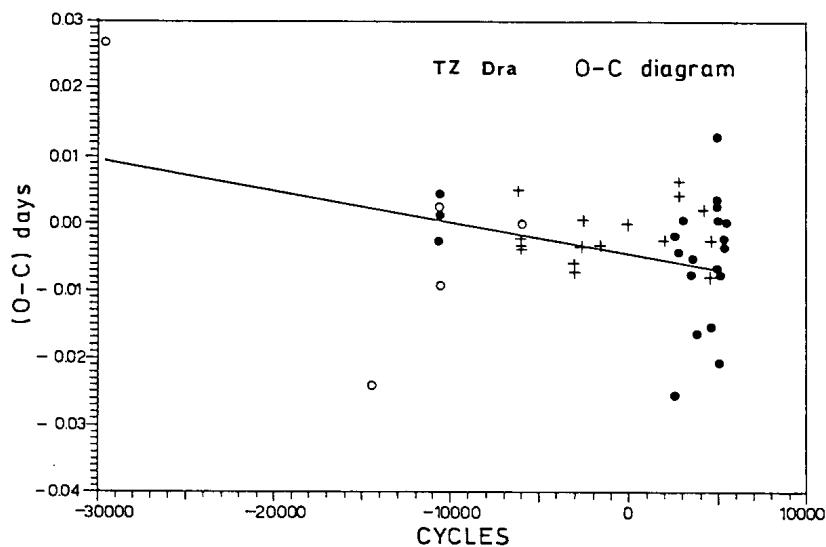


Figure 1 The O-C diagram of TZ Dra based on primary minima times only and according to Kholopov's et al., (1985) ephemeris formula. (Notations: Open circles (○) represent the photographic minima times, crosses (+) the photoelectric and filled circles (●) the visual ones.

and

$$\text{Min I} = 2442966.482 + 0^d.8660347 E \quad (\text{III})$$

(Kholopov et al., 1985)

TABLE I
(New Photoelectric Minima Times of TZ Draconis)

| Hel. JD | $(O-C)_x$ | E_x | $(O-C)_{xx}$ | E_{xx} | $(O-C)_{xxx}$ | E_{xxx} |
|-----------|-----------|---------|--------------|----------|---------------|-----------|
| 2440000.+ | days | | days | | days | |
| 5488.3792 | .0141 | 13414 | .0192 | 8749 | .0042 | 2912 |
| 5495.3097 | .0164 | 13422 | .0214 | 8757 | .0064 | 2920 |
| 5497.4786 | .0202 | 13424.5 | .0252 | 8759.5 | .0102 | 2922.5 |
| 5585.3717 | .0109 | 13526 | .0159 | 8861 | .0008 | 3024 |

From all the minima times of TZ Dra found in the literature, Table II was made, in which the (O-C) values have been calculated using Kholopov's et al. (1985) ephemeris formula. Using the data of Tables I and II, the figure 1 was drawn, which represents the O-C diagram of TZ Draconis, in which crosses denote the photoelectric minima times, open circles the photographic and filled circles the visual ones. From the data of Table I only the three primaries have been taken into account, since the system appears to have an eccentric orbit. As one can see from Figure 1, a linear least square fitting to all the data shows a small change in the period of TZ Dra, but since almost all the minima times (with the exception of the first two and some visual) lay arround zero this must not be true.

TABLE II
Minima Times of TZ Draconis
(According to Kholopov's et al. ephemeris formula)

| Hel. JD 2400000.+ | (O-C)*** days | E*** | Obs. Kind | Reference |
|----------------------|------------------|--------|--------------|-------------------------|
| 17445.332 | 0.027 | -29469 | pg | Plavec, 1964 |
| 30588.272 | -0.024 | -14293 | pg | " |
| 3852.339 | 0.006 | -10524 | pg | " |
| 3864.460 | 0.003 | -10510 | v | " |
| 3865.320 | -0.003 | -10509 | v | " |
| 3871.390 | 0.005 | -10502 | v | " |
| 3884.367 | -0.009 | -10487 | pg | " |
| 3950.196 | 0.001 | -6024 | pg | " |
| 7840.4202 | -0.0024 | -5919 | pe | " |
| 7866.3999 | -0.0038 | -5889 | pe | " |
| 7897.5771 | -0.0038 | -5853 | pe | " |
| 7905.3730 | -0.0022 | -5844 | pe | " |
| 7911.4342 | -0.0033 | -5837 | pe | " |
| 7911.4353 | -0.0022 | -5837 | pe | " |
| 40394.353 | -0.006 | -2970 | pe | Pohl et al., 1970 |
| 0419.4667 | -0.0072 | -2941 | pe | " |
| 0814.3824 | -0.0034 | -2485 | pe | Kizilirmak et al., 1971 |
| 0852.492 | 0.001 | -2441 | pe | " |
| 1519.3351 | -0.0029 | -1671 | pe | Kizilirmak et al., 1974 |
| 2966.4820 | 0.0000 | 0 | pe | Pohl et al., 1977 |
| 4770.4300 | -0.0023 | 2083 | v | BBSAG No. 55, 1981 |
| 5223.3430 | -0.0254 | 2606 | v | BBSAG No. 62, 1982 |
| 5275.3290 | -0.0015 | 2666 | v | BBSAG No. 64, 1983 |

TABLE II (cont.)

| Hel. JD 2400000.0 | (O-C)*** days | E*** | Obs. Kind | Reference |
|----------------------|------------------|------|--------------|--------------------|
| 45478.8447 | -0.0040 | 2901 | pe | Faulkner, 1986 |
| 6335.341 | -0.016 | 3890 | v | BBSAG No. 78, 1985 |
| 6657.523 | 0.002 | 4262 | v | BBSAG No. 81, 1986 |
| 6962.349 | -0.015 | 4614 | v | BBSAG No. 84, 1987 |
| 6981.411 | -0.008 | 4636 | v | " |
| 7000.469 | -0.003 | 4658 | v | BBSAG No. 86, 1988 |
| 7324.372 | 0.003 | 5032 | v | BBSAG No. 88, 1988 |
| 7330.424 | -0.007 | 5039 | v | BBSAG No. 89, 1988 |
| 7362.474 | 0.000 | 5076 | v | " |
| 7369.404 | 0.002 | 5084 | v | " |
| 7382.377 | -0.016 | 5099 | v | " |
| 7467.276 | 0.012 | 5197 | v | BBSAG No. 90, 1989 |
| 7480.247 | -0.008 | 5212 | v | " |
| 7757.383 | -0.003 | 5532 | v | BBSAG No. 92, 1989 |
| 7770.372 | -0.004 | 5547 | v | " |
| 7816.275 | -0.001 | 5600 | v | BBSAG No. 93, 1990 |

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