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THREE NEW VARIABLE STARS

A continuing photographic patrol has resulted in the discovery of three more variable stars brighter than 10th magnitude. The positions and preliminary ranges, types, and periods are presented in Table I, which continues the list in Kaiser et al. (1990).

TABLE I.

Var. designation	RA (1950)	Dec (1950)	Range (v)	Type	P (days)
DHK 14	4 ^h 55 ^m 50 ^s	+24° 25.2'	8.0 - 8.5	EB	1.385
DHK 15	9 51 18	+10 29.7	7.8 - 8.0	SR:	?
DHK 16	9 41 22	+25 35.1	9.2 - 9.9	EA	1.374

DHK $14 = BD + 24^{\circ}719$, HD 31679, SAO 76868 (Tau)

The spectral type is B5. The light curve shows continuous variation, with the primary minimum about 0^m , deep and the secondary 0^m . Photoelectric observations are reported by Williams et al. (1990).

DHK 15 = BD $\pm 10^{\circ}$ 2067, HD 85720, SAO 98835 (Leo)

The spectral type is M7. Variability within the given magnitudes has been confirmed by photoelectric measures, which are continuing and will be reported when the period (if type SR) and full amplitude are better defined.

DHK $16 = BD + 26^{\circ}1996$, SAO 80992 (Leo)

The spectral type is GO. Photoelectric observations show a primary minimum 0.7 deep and secondary about 0.7. The eclipses are three hours in duration. A full report will be published at the end of the observing season.

Some of the information in this report was obtained from the SIMBAD data retrieval system, database of the Strasbourg, France, Astronomical Data Center, for which I thank Joyce Rey-Watson of the Harvard-Smithsonian Center for Astrophysics and Elizabeth Waagen of the AAVSO staff. The variables were detected with a projection blink comparator (PROBLICOM) as described by Mayer (1977). I am grateful to Marvin E. Baldwin, David B. Williams, and James E. Wood for observations to confirm variability and determine periods and types. Their observations will be presented more fully in later reports on each star.

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