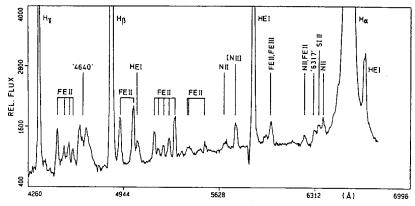
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NOVA OPHIUCHI 1988 MONITORED 2.5 MONTHS AFTER OUTBURST

During the nights of June 21 and 22, 1988 (about 2.5 months after outburst) 12 spectrograms (resolution 4Å) of Nova Ophiuchi 1988 (V 2214 Oph) were obtained using the Boller & Chivens spectrograph and a CCD detector attached to the 1.52m telescope of the European Southern Observatory at La Silla / Chile. The dominant lines in the recorded wavelength region are Balmer emissions (H_{α} to $H_{\gamma})$ and He I $\lambda\,5876.$ H_{α} is extremely strong and is blended with He I and presumedly with NII and [NII] lines. Weaker emissions are mostly due to Fe II but also NII and [NII] are present. All Balmer lines as well as He I λ 5876 exhibit asymmetric profiles with a pronounced red peak whereas the faint lines are more symmetric. The following values of heliocentric radial velocity (uncertainty about 10 km/s), FWHM, FWZI (all in km/s) and equivalent width (Å) were derived for the prominent lines: Ha: -144, 1060, 15 000:, 6500:; H_{β} : -71, 950, 3950, 510; H_{γ} : -28, 985, 3525, 150; He I λ 5876: -167, 1045, 2760, 90. The five strongest Fe II lines yield a radial velocity of -220 ± 17 km/s. The results for He I λ 6678 and [N II] λ 5755 are -247km/s and -294km/s respectively. The absorption feature at the blue side of He I $\lambda\,5876\,$ has a velocity of -1840 km/s.



Average spectrum of Nova Ophiuchi 1988 based on observations obtained on June 21 and 22, 1988. It has been truncated to show the weak emission lines.

A few photoelectric measurements obtained on June 23.029 UT using the single-channel standard photometer at the ESO 50cm telescope gave V=10.915, U-B=-0.335, B-V=+0.664, V-R=+1.822, V-I=+1.115 (uncertainty 0.045 mag in V-I, otherwise less than 0.02 mag).

This note may complement spectroscopic and photometric measurements performed by Jablonski et al. (1988) about one week prior to these observations.

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Reference:

Jablonski, F.J., Kirhakos, S.D., Diaz, M.P., Baptista, R.: IAU Circular No. 4619, 1988