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MORE ABOUT KX TrA = Cn1-2 = PK 326 - 10°1 = He 2-177*

Recently Feibelman (1988) and Allen (1988) commented on the identification and possible outburst properties of the symbiotic star KX TrA. I like to add that it is also listed and (properly) identified in the *Reference Catalogue and Atlas of Galactic Novae* (Duerbeck 1987); it was included because Carlson and Henize (1974) classified it as a slow nova from its spectral appearance, which showed some resemblance to RR Tel. A light curve of the object for the years 1890-1975 was given by Liller (1974). Any major outbursts during this time interval can be excluded.

KX TrA was observed, in 1986, with the B&C Cassegrain spectrograph and IDS detector at the ESO 1.52 m-telescope (Fig. 1). I tend to compare its spectrum with RT Ser, as it appeared in 1986 (Fig. 2), except that in the latter object the [O III] lines are essentially absent. Comparing

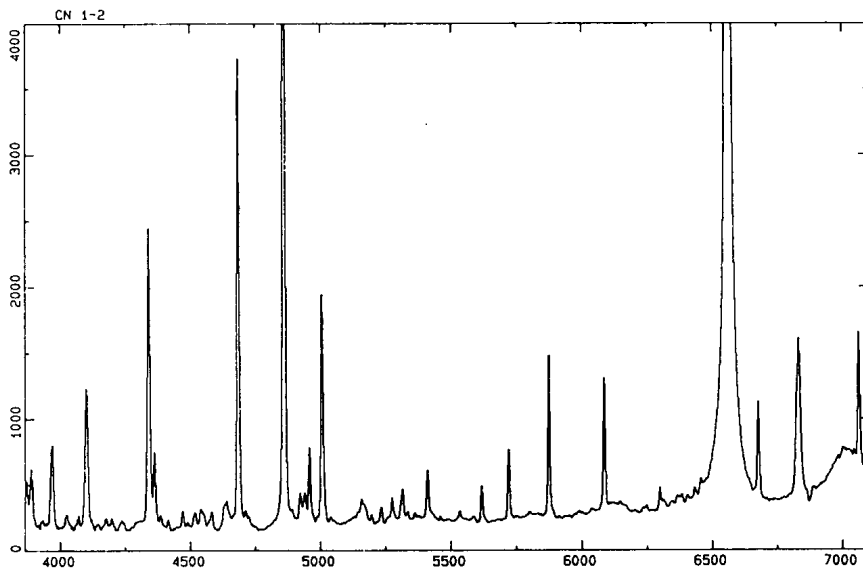


Fig. 1 The spectrum of KX TrA

*Based on observations collected at the European Southern Observatory,
La Silla, Chile

Allen's (1984) spectrum of KX TrA in 1978 with that in 1986, a noticeable relative increase in the strength of He I 4686 is found. It should be noted that the [Ca VII] 561.6, 493.9 lines are relatively strong in both stars. RT Ser also qualifies as a star with a normal M-type companion (Kenyon 1986), a property considered by Allen to be possibly significant for a comparison.

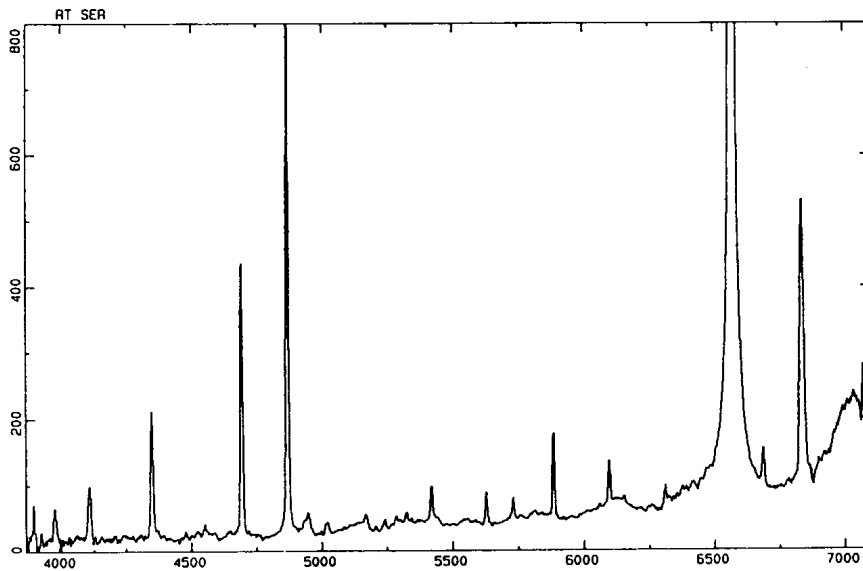


Fig. 2 The spectrum of RT Ser

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