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PHOTOELECTRIC OBSERVATIONS OF ST AQUARII

The eclipsing binary ST Aquarii (BD-7°5753, HD 211965, SAO 146035) was discovered by Miss Leavitt at Harvard (Pickering, 1908). She classified the system to be an Algol type. Zacharov (1928) estimated the period to be 0.390455. Later on Mergentaler (1928) discovered that the period should be much longer than that estimated by Zacharov. Zacharov (1930) reclassified the system to be β -Lyrae type, and he found the period about twice the original value, or 0.7810161. The system consists of a bright A7 component accompanied with a faint G8IV component (Roman, 1956).

The eclipsing binary ST Aqr was observed photoelectrically at Kottamia observatory (Egypt) on 5 nights during October to December 1978. The cassegrain focus of the 74-inch reflector was equipped with a cooled two-channel photometer. The detector was an EMI 9862B/350 photomultiplier tube, powered by an Isotopes Development Ltd. type 1388 D.C. amplifier and then fed to a Honeywell chart recorder. Two wide - band filters (B, V) have been used in this work. The blue filter was a BG12 1mm + GG385 2mm, and the yellow filter was a BG18 1mm + GG495 2mm Schott Glass B filter. The effective wavelength for B and V filters were 4258 Å and 5385 Å, respectively, and the bandwidth for the filters was 982 Å for B and 981 Å for V.

For comparison and check, the stars BD-7° 5751 and BD-7°5755 were used, respectively. A total of 386 and 370 observation points were obtained in B and V filters. The phases of the observations have been computed from the light elements given by Gleim (1973):

Min I (Hel.) = $2441 \ 236.316 + 0.78099525 E$

The light curves for V and B are shown in Figure 1 and 2. Table I shows the times of minima and 0-C values derived from the light curves.

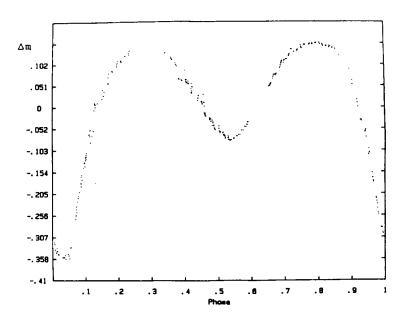


Figure 1

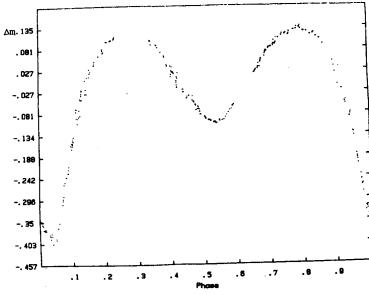


Figure 2

Table I
Times of minima and O-C values

Minimum	Observed (Hel.J.D.)	Calculated (Hel.J.D.)	o-c
I	2443801.9104	2443801.8854	0.0250
II	2443835.8877	2443835.8587	0.0290

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