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RADIAL VELOCITIES OF THE BINARY STAR HD 37020 (91 Ori A)

HD 37020 was discovered as an eclipsing binary by Lohsen (1975). Baldwin (1976) and Franz (1977) estimated its period P=65.4325 days. A tentative spectroscopic orbit was determined by Lohsen (1976) using the radial velocity data available in the literature and Baldwin's period.

Systematic observations of HD 37020 were made at the Brera-Merate Observatory, using the 137 cm Ruths telescope and the B&C spectrograph mod. 31523.

The average of 13 photographic spectra taken in the range 3700-4700 A allowed a re-classification of HD 37020 as a B3 III-IV star. Spectra were also taken in the sub-range 4050-4450 A using a Reticon device (Bonanno et al., 1984); in both cases the reciprocal dispersion was 35A/mm. Nebular lines were adopted as a reference for radial velocities; at last, photographic and Reticon results are perfectly compatible.

Orbital elements (Table I) and a velocity curve (Figure 1) were obtained, assuming the light minimum as phase zero and a period of 65.4325 days. Radial velocities are listed in Table II.

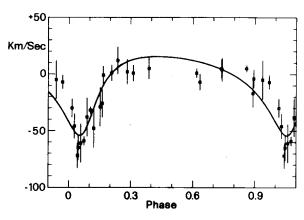


Figure 1

Table I

 $a_1 \sin i = 0.184 \text{ A.U.}$ e = 0.476 $\omega = 201^{\circ}.7$ $V_o = 3.9 \text{ km/sec} + V_n \text{ (%)}$ $T_o = \text{J.D. } 2444201.623$ (%) $\mathbf{V_n}$ is the radial velocity of the Orion Nebula

Table II

Julian Date	Velocity		Error	Julian Date	Velocity		Error
2446819.41	-26	+/-	12	2447184.33	5	+/-	8
822.39	1		7	192.28	5		3
824.28	12		12	194.27	-17		12
827.28	2		11	197.36	- 5		17
829.29	1		7	203.26	-46		11
834.29	5		9	204.27	-72		11
849.32	1		4	205.27	-61		17
850.31	- 7		7	206.30	-59		4
857.31	4		10	207.30	-38		17
867.33	-4		8	208.28	-32		3
872.29	- 7		6	209.30	-48		17
875.30	-30		8	211.31	-29		17
877.32	-65		11	212.29	-1		8

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