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A NEWLY DETECTED RR LYRAE STAR

The star BD+40°0060 = NSV 134(Strohmeier, 1958; Nikulina, 1959) is supposed to be a rapid irregular variable in the New Catalogue of Suspected Variable Stars. A more detailed inspection of the object revealed that it consists of two stars of almost equal brightness, the angular distance between them being about 6 seconds of arc. Photoelectric UBV observations were carried out by the author during 24 nights from 1985 September 11 to 1986 December 5 with the Sonneberg 60 cm telescope II and the Piszkéstető 50 cm telescope (Hungary). Because of the small angular distance of the pair it was not possible to measure the stars separately. The magnitude difference in V to the comparison star BD+40°0056 varies between 0.5 to 1.0 mag. One of the stars, probably the preceding one, has clearly turned out to be an RR Lyrae star of type "ab". The companion would then have to be a late G type star according to the classification by Götz and Wenzel (1962). The observations in V are best reproduced by the following elements:

$$JD_{hel}(Max.) = 244 6764.241 + 0.470568 \cdot E.$$

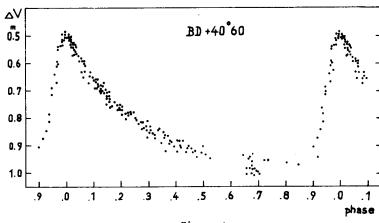


Figure 1

Whether or not the object is a physical pair is still unresolved.

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