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Further photographic UB_V photometry has been obtained for this peculiar, variable emission-line F supergiant from plates taken with the Uppsala Southern Station Schmidt telescope at Siding Spring Observatory, Australia. The results are, with estimated probable errors ± 0.1 mag.:

Date	V	B	U	Observer
20 Jan. 1985	9.5	10.2	10.9	B. Pettersson
21 Jan. 1985	9.6	10.3	10.8	"
14 Feb. 1985	9.2	10.1	10.6	"
10 Mar. 1986	9.3-9.7	10.3	11.0:	P. Magnusson

From 14 Feb. 1985 there are two sets of plates taken two hours apart; the small differences are possibly not significant, as they are only slightly larger than the estimated errors. On 10 March 1986 four sets of plates were taken in direct succession (30-45 min between plates of the same colour). The star was then fainter in V at the beginning and end of the observations, and brighter in between, whereas there seems to have been a small increase in the brightness in U at the end; B remained constant to within the presumed errors. This supports the suggestion of rapid variations put forward by Welin (1981).

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Reference:

Welin, G. 1981, *Inf. Bull. Var. Stars* No. 1940