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UBV PHOTOMETRY FOR BD+37^o 2356

The variable star BD+37^o 2356, classified as W UMa type variable star by Zhukov (1982), was observed in March and May 1984 using UBV photometer attached to the 91 cm reflector at McDonald Observatory. Same comparison star (HD 113 730 = BD+37^o 2360) with Zhukov's was used and about 300 observations in U, B and V bands were taken for the program star.

Six moments of minima were determined, their JD(hel) are as follows:

2445 769.7314 (II), 2445 770.6919 (I), 2445 770.8840 (II),
2445 773.7655 (I), 2445 773.9575 (II), 2445 841.7710 (I)

The new epoch of the primary minimum and the orbital period are:

$JD(hel) = 2445\ 841.7710 (\pm 0.0003) + 0.^d.384\ 2105 (\pm 0.000\ 0001)E$

The new period is a little longer than 0.^d.38416 given by Zhukov.

The light curve of V magnitude is shown in Figure 1. The eclipse depths at primary and secondary minima are almost equal. $V_{max} = 10.^m.228 (\pm 0.^m.008)$, $V_{min} = 10.^m.689 (\pm 0.^m.008)$. The colour indexes, shown in Figure 2 and Figure 3,

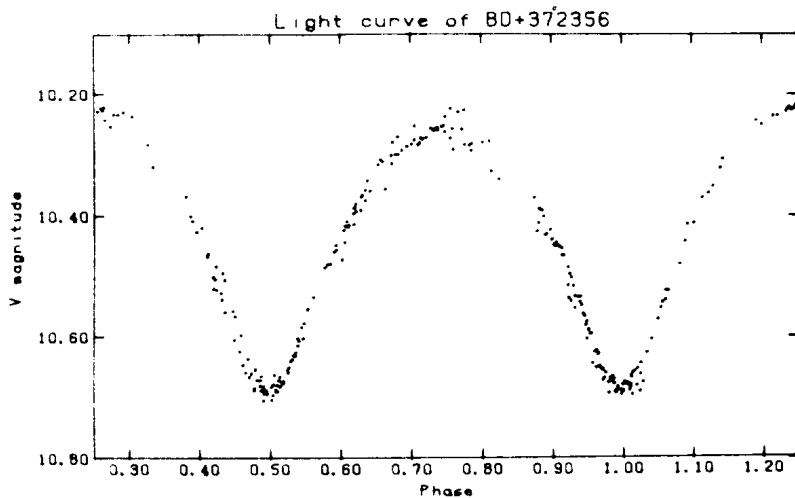


Figure 1

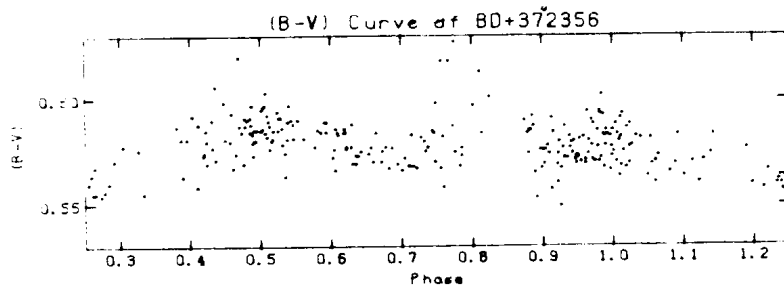


Figure 2

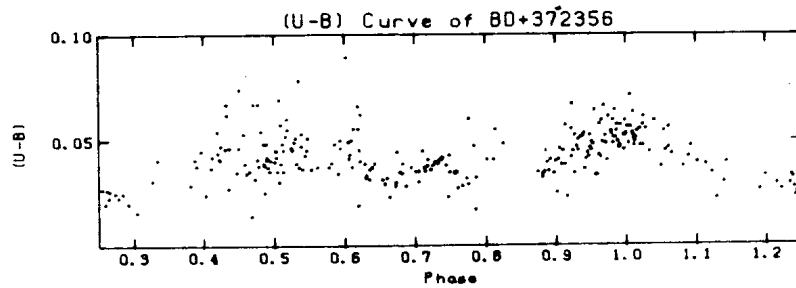


Figure 3

vary lightly with phase. The average colour indexes are:

$$\overline{(B-V)} = 0.^m.578 \quad , \quad \overline{(U-B)} = 0.^m.045 \quad .$$

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Reference:

Zhukov, G.V. 1982, Inf. Bull. Var. Stars, No. 2191.