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V 1050 Aql NOT OF U GEMINORUM BUT OF MIRA TYPE

The following notes exist in the literature on this variable star  
(= S 8194):

- C. Hoffmeister, *Astron. Nachr.*, 288, 49 (1965) - discovery, U Gem type,  
maxima comparatively numerous;
- M. Petit, *Ciel et Terre*, 86, 229 (1970) - U Gem type?, cycle length  $> 80^d$ ;
- N. Vogt and F.M. Bateson, *ESO Scient. Prepr.* 161 (1981) - no additional  
confirmation, map.

When estimating the star on Sonneberg astrographic plates centered at  $\gamma$  Aql  
I noticed that it is a Mira variable with a period of about 291 days; ele-  
ments and light curve will be published in *Mitt. Veränd. Sterne*. On the  
charts of the POSS and therefore on Vogt's and Bateson's map, near the posi-  
tion of the variable three stars can be seen, which are not separated on our  
plates. The south-preceding star of this triple is a red and should be the  
Mira variable. Whether the bluish star marked by Vogt and Bateson at the  
triple's centre as being V 1050 Aql is also variable cannot be proved by our  
plates.

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