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BV PHOTOMETRY OF THE SUPERNOVA 1982 IN NGC 2268

BV photometry of the supernova in NGC 2268 discovered by Wild (1982) is reported.

Table I

BV photometry of the supernova

| Date | UT | B | V |
|---------|---------------------------------|-------|-------|
| 27/2/82 | 21 ^h 17 ^m | | 14.00 |
| 27/2/82 | 22 50 | | 14.05 |
| 01/3/82 | 19 30 | | 14.05 |
| 02/3/82 | 20 05 | | 14.40 |
| 06/3/82 | 21 02 | 15.4 | |
| 06/3/82 | 20 18 | | 14.45 |
| 14/3/82 | 19 43 | 16.00 | |
| 14/3/82 | 20 15 | | 14.65 |
| 15/3/82 | 20 15 | 16.05 | |
| 15/3/82 | 20 41 | | 14.65 |
| 24/3/82 | 20 15 | 16.45 | |
| 24/3/82 | 21 41 | | 15.30 |
| 26/3/82 | 20 10 | 16.50 | |
| 26/3/82 | 21 56 | | 15.35 |
| 14/4/82 | 20 58 | 16.85 | |
| 14/4/82 | 20 15 | | 15.80 |
| 18/4/82 | 21 08 | 16.95 | |
| 14/5/82 | 21 49 | | 16.1) |

The observations were obtained with the 0.4 m Schmidt telescope in Metzlerlen on plates 103a-0, 103a-D and filters Schott 2 mm GG 13 in B and Schott 2 mm GG 11 in V.

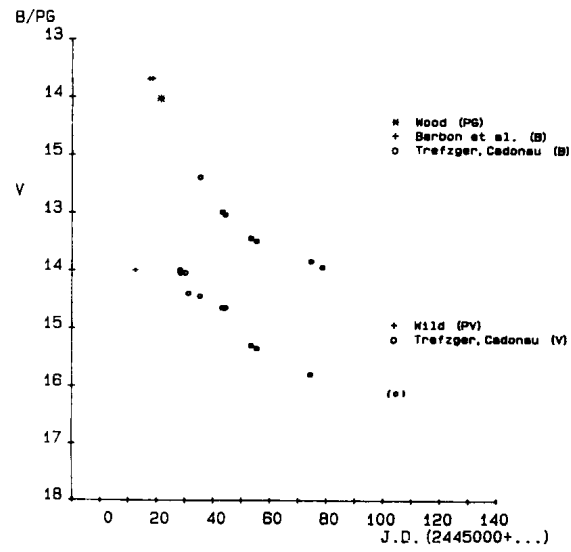


Figure 1: B light curve of SN 1982 Typ I in NGC 2268



Figure 2: SN,NGC 2268 and comparison sequence

Table II
Comparison sequence

| No. | B | V | No. | B | V |
|-----|-------|------|-----|-------|------|
| 1 | 15.4 | 14.5 | 34 | 16.0 | 14.6 |
| 2 | 15.3 | 14.0 | 35 | 16.1 | 15.3 |
| 3 | 14.6 | 14.0 | 36 | 16.0 | 15.1 |
| 4 | 15.9 | 15.0 | 37 | 16.1 | 15.1 |
| 5 | 16.9 | 16.1 | 38 | 16.0 | 15.0 |
| 6 | 17.0 | 16.4 | 39 | 16.0 | 14.8 |
| 7 | 15.7 | 14.7 | 40 | 13.7 | 12.7 |
| 8 | 15.9 | 15.2 | 41 | 16.8 | |
| 9 | 16.3 | 15.6 | 42 | 17.3 | |
| 10 | 16.9 | 16.2 | 43 | 16.4 | |
| 11 | 17.1 | 16.1 | 44 | 17.0 | |
| 12 | 14.7 | 13.6 | 45 | 16.1 | |
| 13 | 14.9 | 13.5 | 46 | 17.2 | |
| 14 | 16.6 | 15.5 | 47 | 14.8 | |
| 15 | 15.2 | 14.3 | 48 | 16.8 | |
| 16 | 16.2 | 15.5 | 49 | 15.7 | |
| 17 | 16.5 | 15.1 | 50 | 17.2) | |
| 18 | 16.8 | 15.8 | 51 | 14.7 | |
| 19 | 16.4 | 15.7 | 52 | 14.5 | |
| 20 | 14.6 | 13.3 | 53 | 14.9 | |
| 21 | 16.3 | 15.5 | 54 | 17.4 | 16.8 |
| 22 | 16.2 | 15.3 | 55 | 16.4 | 15.6 |
| 23 | 16.5 | 15.5 | 56 | 16.5 | 15.8 |
| 24 | 17.2 | 16.3 | 57 | 16.1 | |
| 25 | 16.2 | 15.4 | 58 | 17.5) | |
| 26 | 15.1 | 14.1 | 59 | 17.5 | 16.9 |
| 27 | 15.2 | 14.3 | 60 | 16.7 | 16.0 |
| 28 | 15.1 | 14.1 | 61 | 17.2 | 16.4 |
| 29 | 16.1 | 15.4 | 62 | 16.6 | 15.8 |
| 30 | 17.3 | 16.6 | 63 | 16.5 | 15.6 |
| 31 | 17.8) | 16.5 | 64 | 15.2 | |
| 32 | 16.4 | 15.6 | 65 | 13.4 | |
| 33 | 16.3 | 15.5 | | | |

The comparison sequence was derived from the North Polar Sequence and two fields of stars between the Polar and the comparison sequence.

R. CADONAU
Dr. C. TREFZGER

References:

- Barbon, R. et al.: 1982, IAU Circ., No. 3671
 Wild, P.: 1982, IAU Circ., No. 3667
 Wood, R.: 1982, IAU Circ., No. 3678