

COMMISSION 27 OF THE I. A. U.
INFORMATION BULLETIN ON VARIABLE STARS

Number 2289

Konkoly Observatory
Budapest
1983 March 3
HU ISSN 0374-0676

NEW DELTA-SCUTI-TYPE VARIABLE IN TRIANGULUM

While observing the eclipsing binary V Trianguli: $1^{\text{h}}28^{\text{m}}57^{\text{s}}$, $+30^{\circ}6'27''$ (1950) a nearby star at $1^{\text{h}}29^{\text{m}}15^{\text{s}}$, $+30^{\circ}7'10''$ (1950) was found to be variable. It was observed in the Washington Photometric System colors C, M, and V on eight different nights from October, 1982 to January, 1983. The amplitude of variation was $0^{\text{m}}.12$ in C, $0^{\text{m}}.083$ in M, and $0^{\text{m}}.07$ in V and was roughly sinusoidal. Data for the two nights with the best coverage, JD2445341 and JD2445351, are shown in Figure 1. SAO 54783 was used as a comparison star. Delta magnitude in V (new variable - SAO 54783) is plotted versus heliocentric Julian Date.

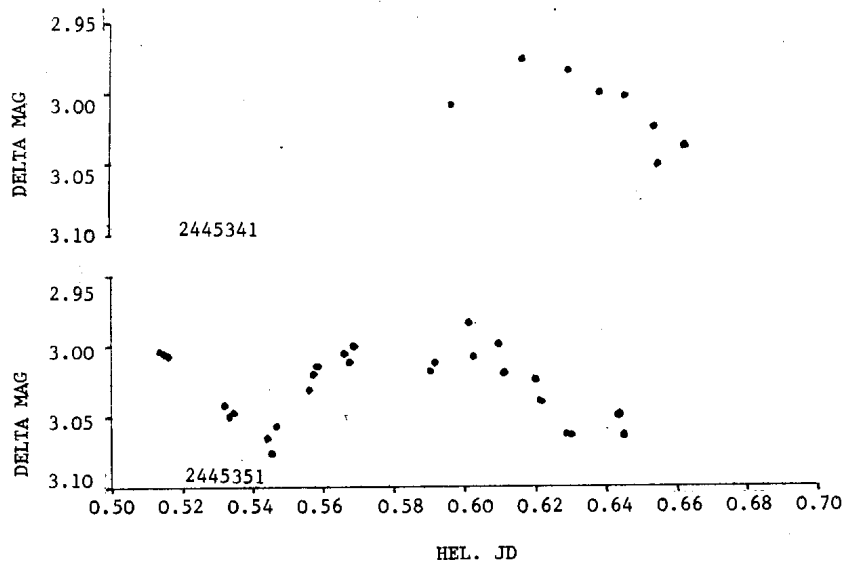


Figure 1

Compared to V Tri at maximum light the new variable is 0.3^m fainter in V, or 11.2^m . Further, it has identical (M-V) and (C-M) indices which implies the same spectral type, A3, as V Tri. A periodogram analysis of the data from all nights indicates a period near 0.10 days, with the possibility of multiple periods. The light amplitude variation, period and spectral type indicate the new variable is a delta-Scuti-type variable.

J. S. SHAW, D. A. FRAQUELLI, D. H.
MARTINS, and D. E. STOOKSBURY
University of Georgia Observatory
University of Georgia
Athens, Georgia 30602