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PHOTOELECTRIC EPOCHS OF MINIMUM LIGHT, UZ PUPPIS

B,V Light curves of the southern eclipsing binary system UZ Puppis (BD-13°2170) were observed photoelectrically in January 1981 with the 50cm telescope at Cerro Tololo InterAmerican Observatory. Two nearby stars were monitored as comparison and check stars. The observations define one primary and one secondary eclipse curve. An iterative procedure based on the Hertzsprung (1928) method was used to determine the epochs of minimum light,

JD Hel. Min. I = 2444613.6983JD Hel. Min. II = 2444615.6871.

These and the photoelectric epochs of minimum light observed by Bloomer (1973) were entered into a least squares solution which yielded the ephemeris

The residuals or O-C's are negative for the primary minima and positive for the secondary minima. However, the data are meager.

A least squares solution utilizing only the times of the primary minima yielded $% \left(1\right) =\left(1\right) \left(1\right) +\left(1\right) \left(1\right) \left(1\right) +\left(1\right) \left(1\right)$

Hel. JD Min. I = 2444613.69829 +
$$0^{\frac{1}{2}}$$
79485120 E.

An orbital analysis of the system is in progress.

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