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VARIABILITY OF BD +4004145 A MEMBER OF NGC 6910

UBV photometric standards in galactic cluster NGC 6910 were observed with the 60 cm reflector at Ostrowik Station of Warsaw University Observatory on February 16, 1982. The unrefrigerated one channel photometer with 27" diaphragm was used. Stars Nos. 1, 2,3,4,5,6,8 (numeration according to Hoag et al., 1961) were investigated through V and B filters. It was found that the star No. $5 = BD + 40^{\circ}4145$ has luminosity $V = 8^{\circ}69$ and colour $B-V=0^{\circ}83$ instead of the catalogue values $V = 10^{\circ}05$ and V = 0.19 (Hoag et al., 1961). Short time scale variations with the amplitude of about $0^{\circ}20$ were observed in both filters.

Table I BV observations of BD +40⁰4145

	Date		JD Hel. 244+	V	B-V
Sep.	21,	1981	4869.2870	10 ^m 03	0 ^m 20
Feb.	16,	1982	5017.6545 .6615	8.41 8.61	1.05 0.89
			.6656	8.59	0.72
			.6689 .6733	8.75 9.05	0.82 0.65
			.6775	8.61:	0.88:
Mar.	27.	1982	.6823 5056.6053	8.79: 8.64	0.76: 0.89
Apr.	14,	1982	5076.5236	8.73	0.81
May. Jun.	5, 4,	1982 1982	5095.4546 5125.5031	10.05 10.07	0.19 0.23
Jun.	25,	1982	5146.3777	10.06	0.18
Jul.	•	1982 1982	5158.4071 5159.4226	9.97 10.02	0.27 0.22

Untypical V and B-V values were also found on March 27, 1982 and April 14, 1982. However, on May 5, 1982 and later the star was in its normal state. We have also reexamined our observations from September 1981. All observations from September 1981 until July 1982 are listed in Table I. Mean errors are 0.015 and 0.020 in V and B-V, respectively. Instrumental quantities were transformed to BV system using standard methods. As far as we know BD +40.04145 has not been known as a variable star.

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Reference:

Hoag, A.A., Johnson, H.L., Iriarte, B., Mitchell, R.I., Hallam, K.L., Sharpless, S. 1961, Photometry of stars in galactic cluster fields, Publ. U.S. Nav. Obs., 17, Part VII, p. 454