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VARIABLE STARS IN THE PLEIADES CLUSTER II

From 1981 October 18 till November 17 we performed photometry on 16 K-type stars in the Pleiades cluster. All of these stars were known or suspected to be variable on the basis of earlier work. Five of the stars were also measured a year before by Alphenaar and Van Leeuwen (1981, hereafter referred to as Paper I). The other 11 stars formed a new selection. All stars were measured using the Walraven photometer and the Dutch 91 cm telescope at La Silla, ESO (Lub, 1979). Positions and search maps of the previously measured stars can be found in Paper I, those for the new selection are given in Fig. 1. The diameter of each field is 12 arcmin and the magnitudes indicated are photographic magnitudes. The limiting magnitude is around  $m_{pg} = 14.5$ . Table I lists the measured stars and their visual magnitudes. The star numbers are from Hertzsprung (1947).

Ten of the eleven newly selected stars were found to be variable and for the other, Hz883, no conclusions could be drawn. For 7 of these stars periods and lightcurves could be obtained, as well as for the 5 remeasured stars. All lightcurves and periods are presented in Fig. 2a, 2b, 2c and 2d as differences in  $10\log(I)$  for the V channel between the given star and the substandard for the cluster, Hz804. The zeropoints for the phase calculations are at  $JD=2444896.5$ .

The periods of the five previously measured stars did not need to be changed. For Hz879 the shorter period given in Paper I turned out to be the right one. The lightcurve of Hz1883 was repeated accurately, but for all others there seem to be disturbing sources active.

The 7 new stars for which periods and lightcurves were obtained are also of the BY Dra type. Four of them, viz. Hz882, 1531, 2034 and 3163, resemble the lightcurve of Hz1883. Hz625 shows a broader minimum than

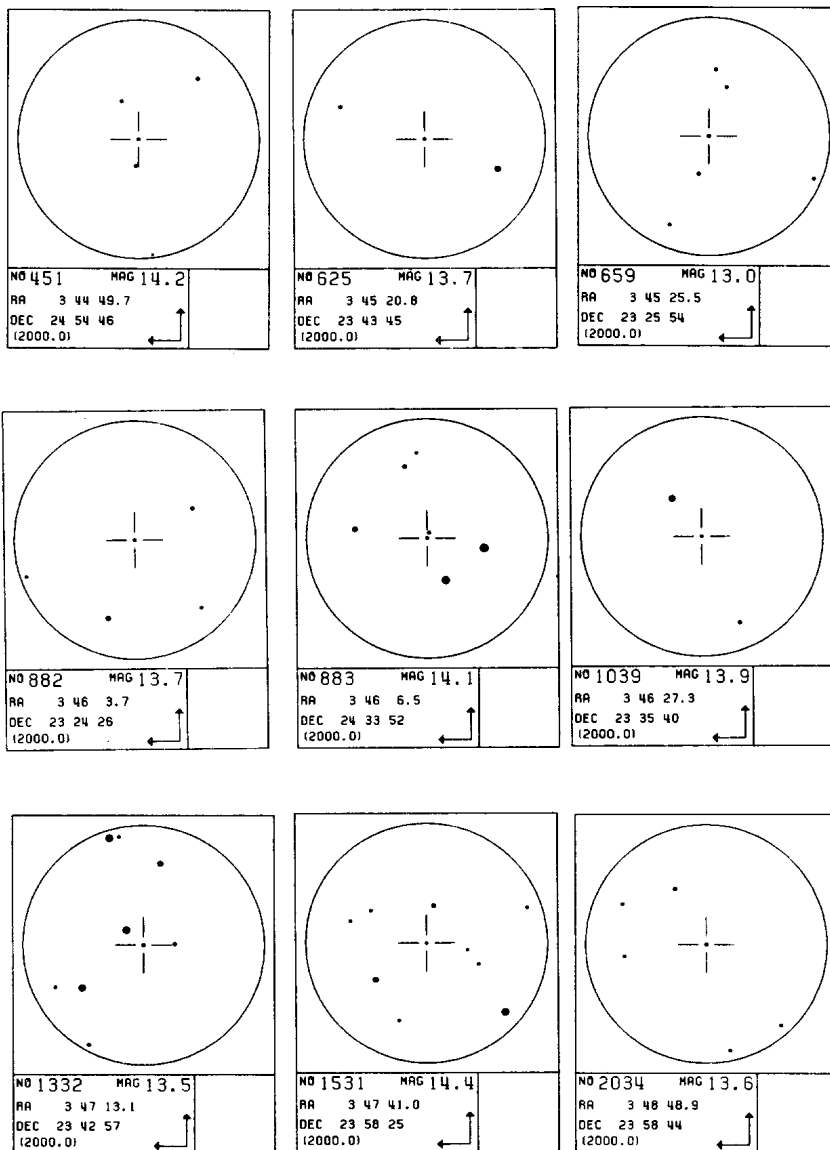


Figure 1

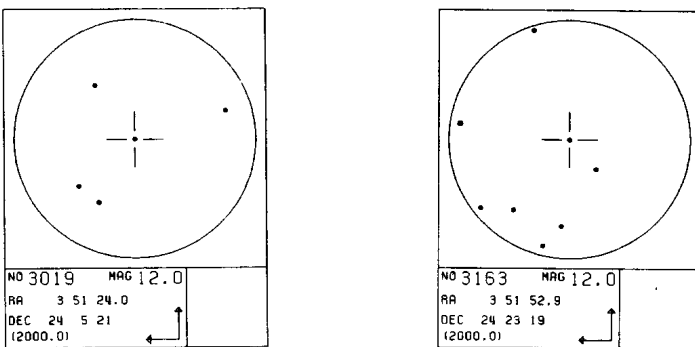


Figure 1 (cont.)

the other stars. The small amplitudes of the stars Hz1039 and 1332 do not permit detailed interpretation. Five of the new selection of stars, viz. Hz451, Hz883, Hz1039, Hz1531 and Hz2034, are known as flare stars (Haro, 1976).

Table I, Measured Stars

Hz	$m_V$	(V-B)
34	12.06	.43
451	13.43	.57
625	12.66	.53
659	12.02	.41
686	13.44	.52
879	12.83	.50
882	12.90	.49
883	13.05	.53
1039	13.05	.57
1124	12.32	.44
1332	12.52	.47
1531	13.58	.56
1883	12.61	.48
2034	12.65	.45
3019	13.49	.57
3163	12.73	.46

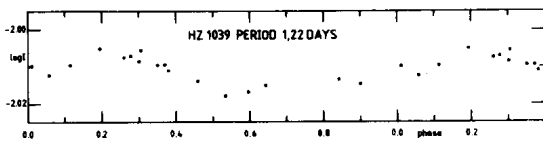
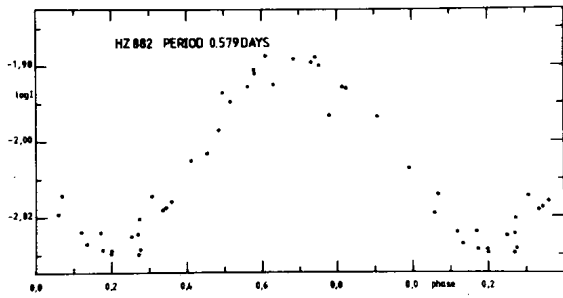
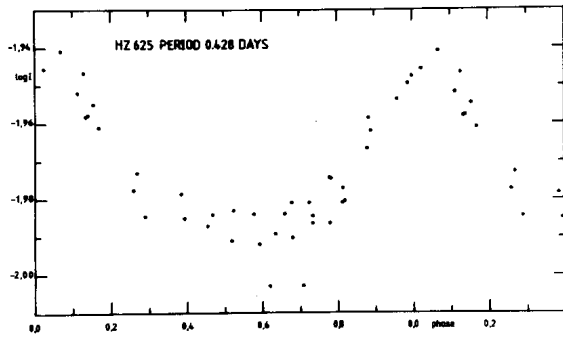


Figure 2.a

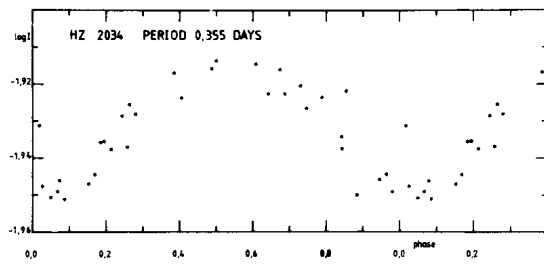
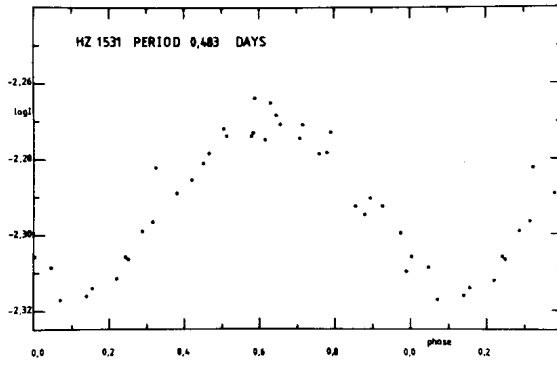
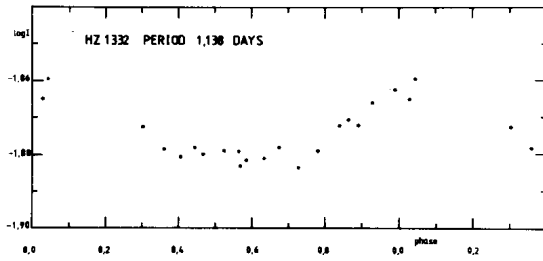


Figure 2.b

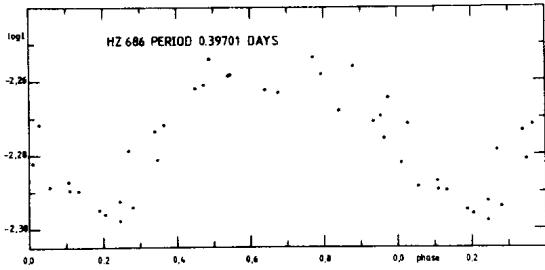
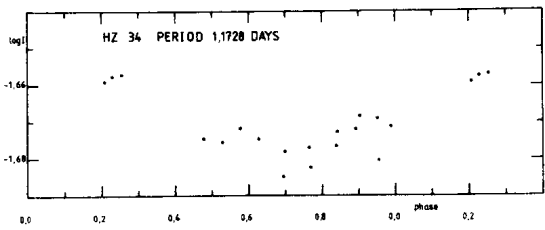
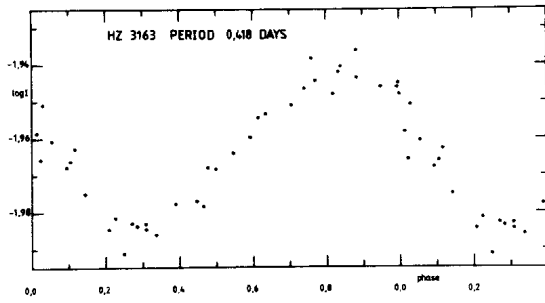


Figure 2.c

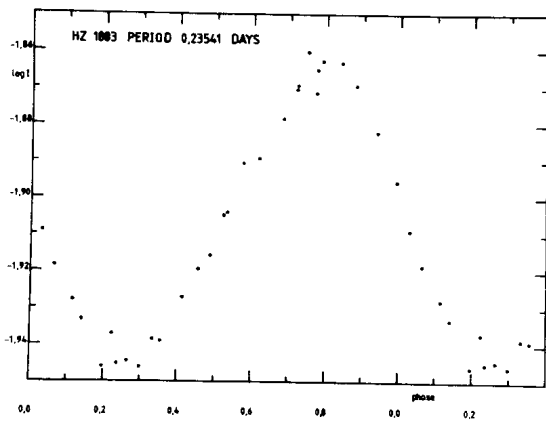
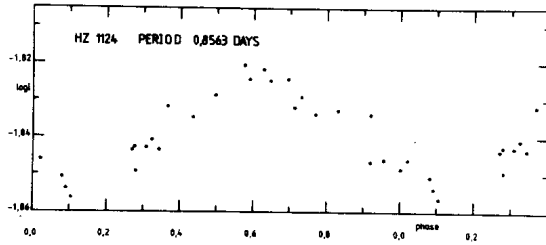
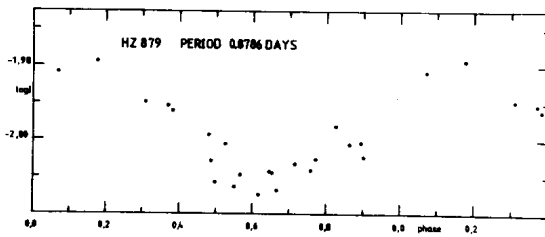


Figure 2.d

In a recent letter to Astronomy and Astrophysics the observations presented here and in Paper I are discussed in more detail, including also some spectroscopic data.

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