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PHOTOELECTRIC LIGHT CURVE OF 44 BOOTIS

A complete light curve of the W UMa-type eclipsing binary 44 Boo was obtained on the night of 8-9 May, 1982. The observations were made on the No. 3 40 cm telescope at Kitt Peak National Observatory. A dry-ice cooled 1P21 photomultiplier and pulse - counting system were used to make the observations in the V band of the UBV system. Observations were made relative to the comparison star 47 Boo.

Observations of the close pair 27/28 LMi provided data for transformation to the standard UBV system. Analysis of the comparison star magnitude versus air mass provided a measure of extinction. Observational times were calibrated with respect to radio station WWV to within 0.1 seconds. All integrations were of 10 seconds duration, and the observations were made in the following order: 6 of the comparison star; 3 of the comparison star background; 60 of the variable star; and 3 of the variable star background; etc. A neutral density filter was used to limit the maximum counts to approximately 200,000 for the 10-second integrations.

The data were reduced at the Fairborn Observatory using the BASIC program described by Genet (1980). This included correction for differential extinction, transformation to the standard system and heliocentric correction of the observational times. Approximately 1500 10-second observations of the variable star were made in about seven hours. These data were grouped and mean counts for each seven observations were used in the analysis. The phase was calculated using the ephemeris of van't Veer (1972).

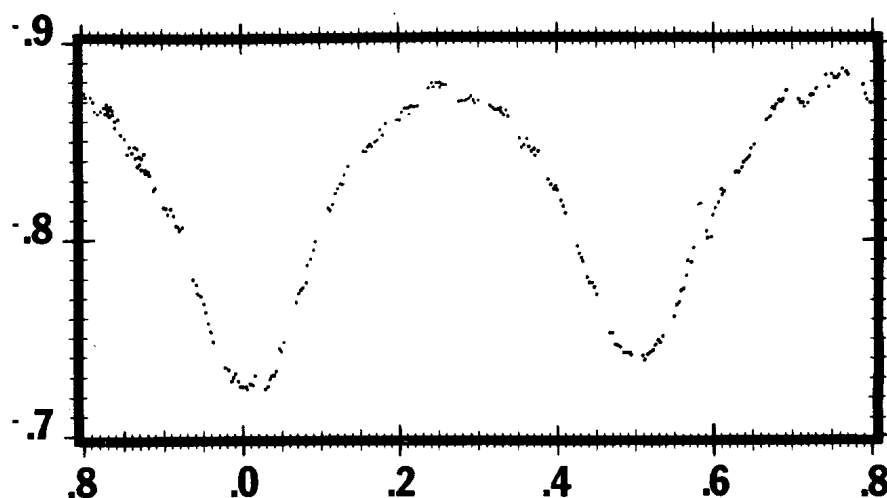


Figure 1

$$\text{Pr. Min.} = \text{Hel. J.D. } 2439370.4222 + 0.2678160 E$$

The resultant light curve is shown in the figure. The data are available as IAU 27 File No. 89. The persistence of certain irregularities in the light curve are being examined with respect to observations made on nights prior to and subsequent to the night's observations reported herein.

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#### References:

Genet, R. M.: 1980 IAPPP Communications, 2, 23.  
 van't Veer, F.: 1972, Astron. & Astrophys., 20, 131.