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Budapest
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SPIKE FLARES OF UV Ceti

During 2-9 December 1975 the flare star UV Ceti [R.A. = $01^{\text{h}}36^{\text{m}}24^{\text{s}}$, declination = $-18^{\circ}13'$ (1950), visual magnitude 12,9 at minimum, spectral type dM5,5e] was monitored at Boyden Observatory with the 41 cm Nishimura reflector. A Johnson B filter and an uncooled EMI 6256 A photomultiplier tube were used in the observations. The two flares of $\frac{I_{\text{0+f}} - I_0}{I_0}$ equal to 14,20 and 3,09 were visually confirmed on the night of 2/3 December in the 15cm finder.

Of the 38 other flares recorded during this period 36 were spike flares, either single flares or in the form of a spike complex. Their $\frac{I_{\text{0+f}} - I_0}{I_0}$ values were between 1,00 and 2,17 with durations generally between 3 and 12 seconds.

The data is presented in the table, following the proposals to flare star observers (1). Examples of light curves representing a selection of the flares are shown in Figures 1-16.

On past occasions we have reported similar spike flares from UV Ceti (references 2-5), using the terminology suggested by Moffett (6).

Our data reduction has been greatly facilitated by the kindness of Prof. F.D.I. Hodgson, Director of the Institute of Groundwater Studies at the Universtiyy, in making available his institutes' Hewlett Packard Model 9825 microcomputer and digitizer.

A.H. JARRETT and J. VAN ROOYEN
Boyden Observatory
Astronomy Department
University of The Orange Free State
Bloemfontein
Republic of South Africa

DATE 1975	MONITORING U.T.	TIME U.T.	FLARE MAXIMUM U.T.	$\frac{I_o + f - I_o}{I_o}$	$2.5 \log \left[\frac{I_o + f - I_o}{I_o} \right]$	$\sigma(\text{mag})$ $= -2.5 \log \left[\frac{ G }{I_o} \right]$	$m_{1.1m} - m_o$ $= \sigma(\text{mag})^{-1} \cdot 1.9$
2 December	18 05 ^m 29 ^s - 18 27 ^m 06 ^s	*	19h37m17s	1,14	0,15	2,10	0,91
	18 31 24 - 18 53 12	*	19 56 13	1,18	0,18	1,90	0,71
	18 57 12 - 19 05 06	*	20 00 42	1,22	0,21	1,90	0,71
	19 07 54 - 22 00 00	*	20 01 12	1,14	0,14	1,90	0,71
		*	20 39 16	1,29	0,27	2,13	0,95
		x	21 20 30	14,20	2,88	2,04	0,85
		x	21 45 07	3,09	1,23	1,98	0,79
3 December	18 01 18 - 18 45 28	*	18 29 25	1,13	0,13	1,98	0,79
	18 49 32 - 19 01 00	**	19 25 52	1,23	0,23	2,03	0,84
	19 04 12 - 22 00 15	*	19 28 55	1,37	0,34	2,03	0,84
		*	19 50 05	1,57	0,49	1,87	0,68
		*	20 39 42	1,39	0,36	2,03	0,84
		*	20 44 43	1,63	0,53	2,03	0,84
		*	20 49 43	1,42	0,38	2,03	0,84
		**	20 56 18	1,49	0,43	2,03	0,84
		**	21 16 45	1,07	0,07	2,03	0,84
		*	21 31 58	1,48	0,43	1,93	0,74
		*	21 35 44	1,55	0,48	1,93	0,74
		**	21 38 41	1,03	0,03	1,93	0,74
		*	21 42 02	1,29	0,28	1,93	0,74

x Visually Confirmed

*

Spike

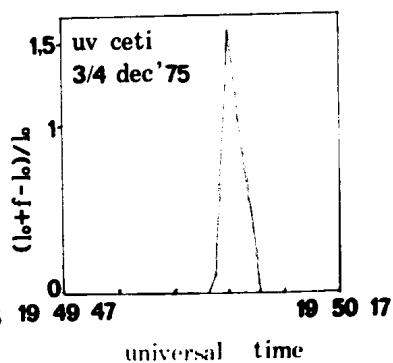
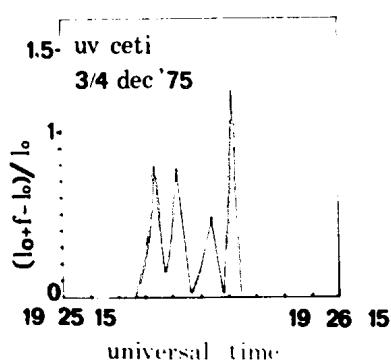
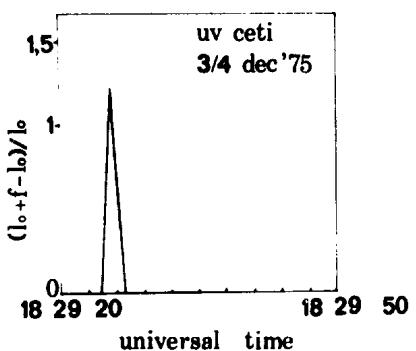
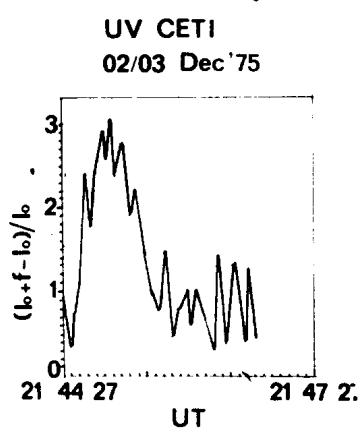
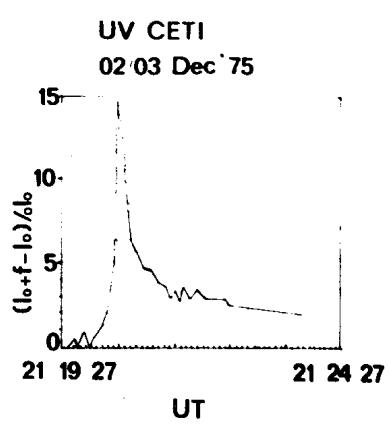
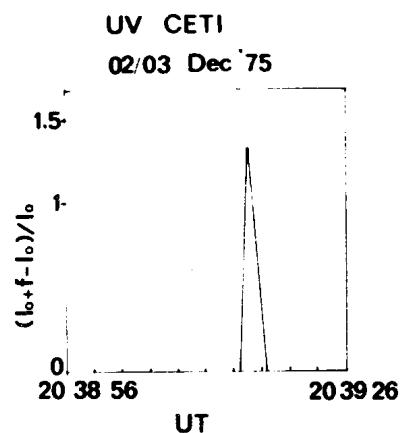
** Spike Complex

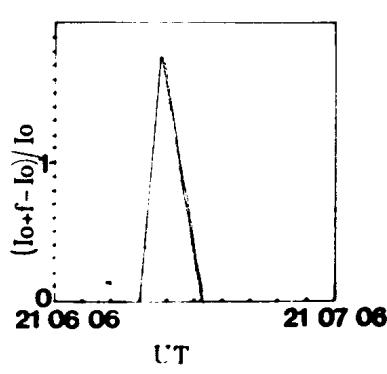
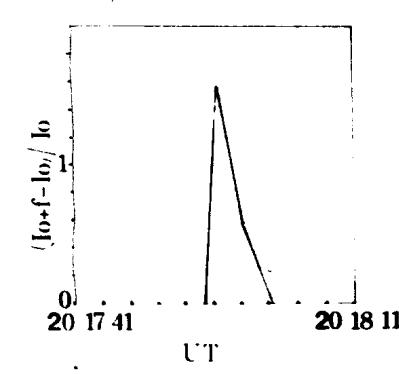
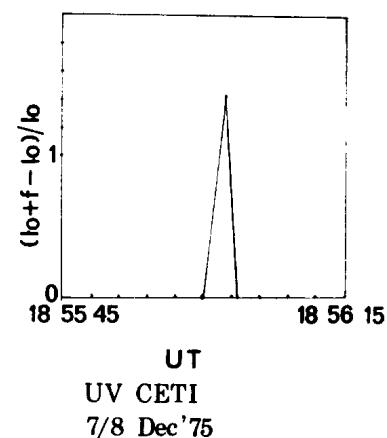
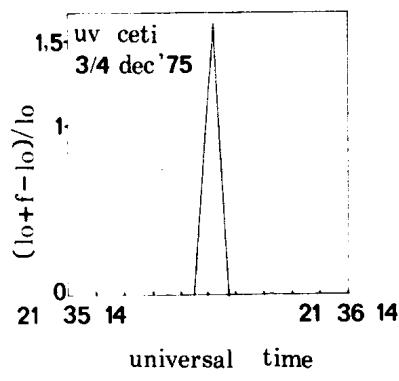
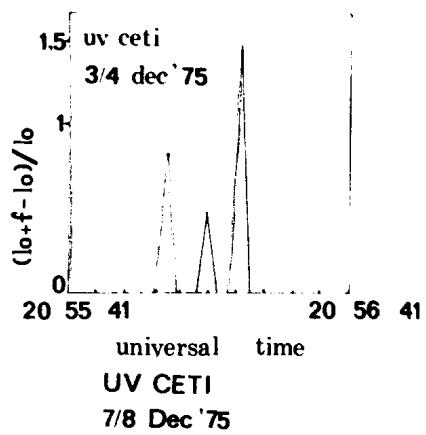
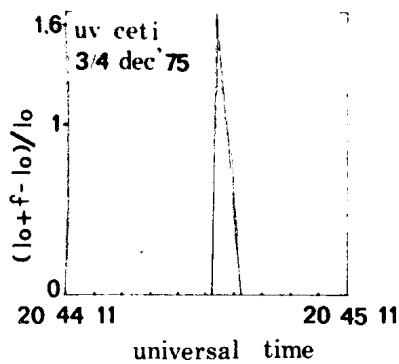
DATE 1975	MONITORING U.T.	TIME U.T.	FLARE MAXIMUM U.T.	$\frac{I_{\text{off}} - I_0}{I_0}$	$2.5 \log \left[\frac{I_{\text{off}} - I_0}{I_0} \right]$	$\sigma(\text{mag})$	$\frac{ O }{I_0}$	$m_{\text{lim}} - m_0$ $= \sigma(\text{mag})^{-1.19}$
7 December	18 ^h 17 ^m 15 ^s - 18 ^h 45 ^m 04 ^s 18 49 02 - 18 36 01 18 39 06 - 22 01 14	*	18 ^h 56 ^m 03 ^s ** 19 58 52 20 17 58 ** 20 22 26 * 20 33 45 * 21 02 46 * 21 06 32 * 21 12 46 ** 21 22 45 ** 21 22 55 * 21 23 38 * 21 24 18 * 21 45 30 * 21 49 20	1,38 1,38 1,56 1,26 1,42 1,28 1,84 2,17 2,04 1,79 1,53 1,28 1,82 1,45	0,35 0,35 0,48 0,25 0,38 0,27 0,66 0,84 0,78 0,63 0,46 0,27 0,65 0,40	1,78 1,78 1,81 1,81 1,81 1,68 1,68 1,68 1,68 1,68 1,68 1,68 1,68 1,68		0,59 0,59 0,62 0,62 0,62 0,49 0,49 0,49 0,49 0,49 0,49 0,49 0,49 0,49
8 December	18 24 26 - 18 34 18 18 38 00 - 18 41 18 18 45 06 - 22 00 18	*	20 55 16 * 21 01 08 * 21 02 38 * 21 39 12 * 21 48 09 * 21 53 27	1,13 1,18 1,00 1,44 1,50 1,39	0,13 0,18 0,00 0,40 0,44 0,36	2,00 1,81 1,81 1,81 1,81 1,81		0,81 0,62 0,62 0,62 0,62 0,62
9 December	18 04 14 - 18 36 14 18 42 00 - 19 07 36 19 12 18 - 23 00 10	NO FLARES RECORDED						0,62

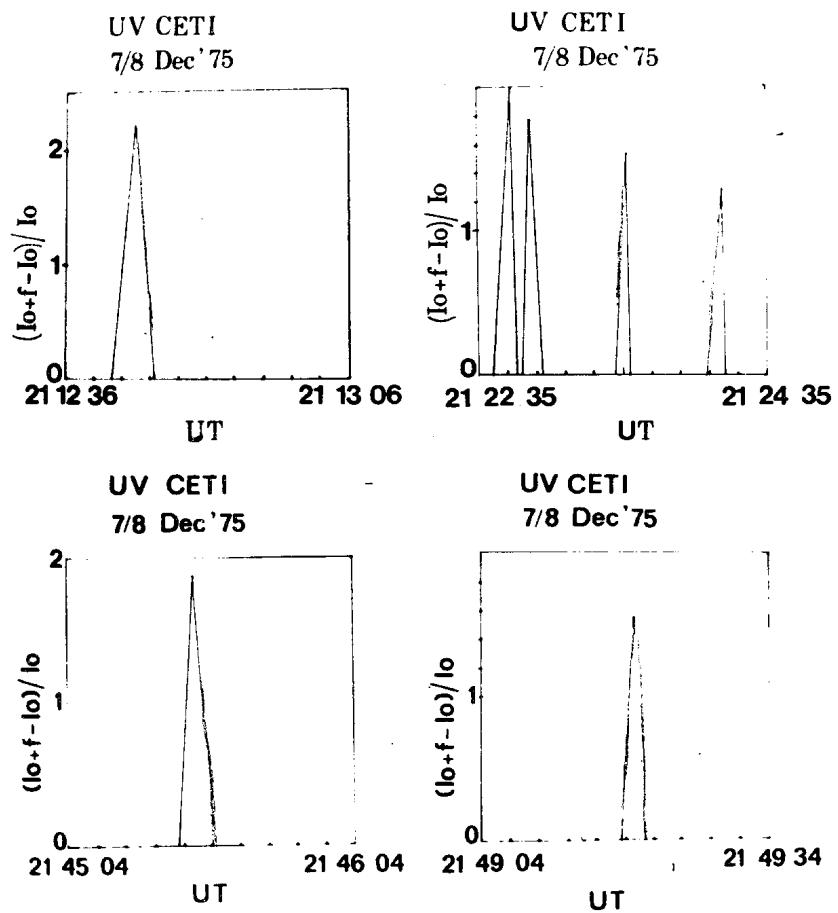
X Visually Confirmed

* Spike

** Spike Complex







References:

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- 5) Jarrett, A.H. and J. van Rooyen, 1979, I.B.V.S. No. 1585
- 6) Moffett, T.J., 1974, Sky and Telescope, 48, 94