

COMMISSION 27 OF THE I. A. U.
INFORMATION BULLETIN ON VARIABLE STARS

Number 1469

Konkoly Observatory
Budapest
1978 September 7

CD -30^o5135

New photographic magnitude determinations have revealed brightness variations of hitherto unknown amplitude in this interesting emission-line F supergiant. The plates were taken by Dr. S. Wrandemark with the Schmidt telescope of the Uppsala Southern Station at Mount Stromlo Observatory, Australia, and the photometric calibration was made by the same means as in my previous report on the star, IBVS No. 1313, 1977. The following magnitudes were obtained (estimated error ± 0.1 mag.):

	V	B	U
Jan. 11, 1978	8.75	9.9	10.7
Jan. 14, 1978	9.1	10.1	10.6

When this is compared to the photometry by Humphreys (PASP 87, 933, 1975), it is seen that the amplitude of the light variations is almost one magnitude in V, and that the star seems to become redder as it brightens.

It has been pointed out earlier (Welin, IBVS No. 1139, 1976) that the star is not a spectroscopic binary, as was suggested by Humphreys (op. cit.). There still remains, however, to relate the brightness variations to variations in the emission and other spectral features, as well as in radial velocity. Simultaneous monitoring of the star by spectroscopic and photometric means for some time seems highly desirable.

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