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**UV Cet**

A continual photoelectric monitoring of the flare star UV Cet was done with the 91 cm reflector of the Okayama Station from 22 September to 5 October 1970.

During the 8.5 hours of monitoring in the magnitude system B, 9 flares were observed as shown in the following:

Date 1970	Time of Monitoring(UT)	Time max.(UT)	Flares $\Delta m(B)$	Dura- tion P	mag
Sept.					
26	17h29m-19h27m	17h57m.3	1.13 <sup>mag</sup>	0.5 min	1 min 0.07 <sup>mag</sup>
27	13 24 -15 19				
	17 12 -17 33	17 19.5	0.53	0.2	0.4 0.15
		17 21.7	0.66	0.2	0.2 0.15
Oct.					
1	18 28 -19 08	18 52.6	2.25	>9.0 >8	0.11
5	15 22 -19 00	15 44.6	0.68	0.1 0.1	0.15
		17 30.2	0.88	0.2 1.5	0.16
		18 00.5	0.70	0.7 3.5	0.19
		18 25.4	0.84	0.3 1.5	0.22
		18 50.6	0.93	1.0 4.0	0.15

$$\Delta m(B) = 2.5 \log (I_{o+f, \text{max}} / I_o)$$

$$P = \int (I_{o+f} - I_o) / I_o dt; \quad C(\text{mag}) = 2.5 \log (I_o + C) / I_o$$

The values of  $\Delta m(B)$ , P and C(mag) are all referred to the total luminosity of L726-8 A+B. and not to L726-8B (UV Cet) only.

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