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## ANOTHER VARIABLE WITH CHANGING PERIOD

Magnitudes for one of my hitherto unpublished variable stars in Sagittarius, at 18h33m07s - 21°14!2 (1900) were estimated by Miss Loretta J.Locicero on some 250 Harvard and 500 Nantucket plates. For the Harvard plates of the MF series (10-inch Metcalf, scale 167"/mm) for the years 1924-34, she found a reciprocal period of 2-467859, and for the Nantucket plates (7.5-inch Cooke, scale 248"/mm) a slightly shorter value, 2.467825. Harvard plates of other series for the intermediate years 1942-51 are sparsely distributed and do not fit either of these periods well. In Figure 1, therefore, I have plotted the phases of ascending light, on the basis of the reciprocal period, 2.467825 and for magnitudes brighter than 13,7, against the Julian day from J.D.2424000 to beyond 2440000. This indicates that the period in changing progressively. All of the observations are then satisfactorily represented (Figure 2) by

Phase = 2.467825 (JD) - n -0.130\*10<sup>-8</sup>(JD -2438500)<sup>2</sup>.

The final term in this expression is represented by the smooth curve through the points in Figure 1. In terms of the period itself we get

Max = JD 2438563.748 + 0\frac{9}{4}052151 n + 0.865.10^{-10}n^2.

The period and change of period are given by

$$P = P_0 (1 + 5.27 \cdot 10^{-10} n)$$

representing a progressive lengthening of the period by 1.85.10 seconds per cycle.

The observations seemed to indicate a possible secular change in amplitude or magnitude at maximum. This can, however, be interpreted merely as systematic errors corresponding to the use of different instruments, different emulsions, and different effective exposures in a field where overlapping images are prevalent. The horizontal spread in the ascending branch of the light curve appears to be random and does not suggest a correction either to the basic period or the secular term.

The insert in Figure 2 is a finding chart showing a field of approximately 10' x 10'.

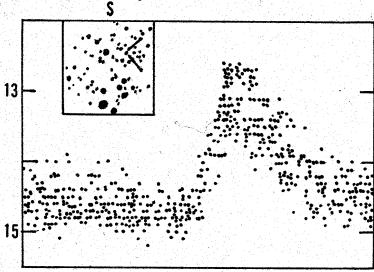


Figure 2

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