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PHOTOGRAPHIC OBSERVATIONS OF UV Ceti (October, 3-18, 1969)

According to the program published in the Inf.Bull. IAU Comm. 27 (1) photographic observations of UV Ceti have been carried out at Abastumani Astrophysical Observatory from 3- to 18 October 1969. Through the use of a special automatic camera adjusted at the primary focus of the telescope almost continuous photography of the star was performed.

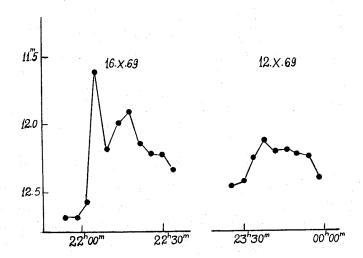
The photographs were obtained on KH-3 film without a filter, the exposures, in most cases, being 4 minutes. Time intervals are given in the table below, when continuous observations were carried out over 9 nights.

8.10.69	$19^{h}53^{m} - 21^{h}48^{m}$	13.10.69	13.10.69 00h00m-00h24m	
0.10.00	22 06 - 22 46	14.10	19 00 -24 00	
9.10	19 10 - 20 22	15.10	00 00 -01 04	
10.10	18 58 - 24 00		18 53 -24 00	
11.10	00 00 - 01 04	16.10	00 00 -01 00	
	19 2 8 - 21 32		19 00 -24 00	
	21 54 - 22 12	17.10	00 00 -00 58	
12.10	19 00 - 20 00		19 22 -24 00	
	20 09 - 24 00	18.10	00 00 -00 34	

The total length of continuous observations amounts to 40^h51^m. More than 600 photonegatives are obtained. A flash at 23^h38^m (U.T.),12 October with an amplitude about 0.3 magn. and a flare at 22^h06^m, 16 October with an amplitude of about 1 magn. have been fixed.

Mean moments of exposure are indicated here, for the cases of greatest brightness. Microphotometry of the photographs, obtained at the moments near to flares, has been done.

The comparison stars were used from (2). The results of microphotometric estimations of magnitudes are given in table and on the diagram (see figure).



16.10.1969			12.10,1969		
No	U.T.	$\mathbf{m}_{\mathbf{p}\mathbf{v}}$	No	U.T.	$\mathbf{m}_{\mathbf{p}\mathbf{v}}$
1	21h54m	12m68	1	23 ^h 26 ^m	12 ^m 46
2	21 58	12.69	2	23 30	12.44
3	22 02	12,58	3	23 34	12.26
4	22 06	11.61	4	23 38	12.14
5	22 10	12,18	5	23 42	12.21
6	22 14	11.99	6	23 46	12.20
7	22 18	11,91	7	23 50	12.24
8	22 22	12.14	. 8	23 54	12.26
9	22 26	12.22	9	23 58	12.43
10	22 30	12.23		•	
11	22 34	12.35	•		

Mean square error of one estimation equals to $\pm 0^{m}05$.

References

- A.D.Andrews, P.F.Chugainov, V.S.Oskanian Comm.27, IAU Inf.Bull.Var.Stars, 1968, No.318
 W.J.Luyten, AphJ 109, 532, 1949.

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