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SPECTRUM OF THE EARLY-TYPE FLARE STAR BD +13°1048

The variability of the sixth magnitude star BD +31 $^{\rm O}$ 1048 was discovered by Andrews 1964 who found an increase of at least 3 $^{\rm m}$ on March 1, 1964, and rapid changes in brightness on March 10 and 14, with amplitudes up to 2 magnitudes. Two spectrograms in the blue region taken at Herstmonceaux and Edinburgh show normal B8 features, in agreement with the HD classification but not in accord with the present scheme of flare stars. However, according to the strict temporal meaning of the word, the object is claimed to be a flare star (Andrews, 1965). The author of the discovery hopes that possibly ${\rm H}_{\alpha}$ could be detected in emission, but this is not the case. Two spectra obtained at Asiago on November $22^{\rm d}3^{\rm h}5^{\rm m}$ U.T., 1964 and November $8^{\rm d}22^{\rm h}53^{\rm m}$ U.T., 1965, both on Kodak 103aF with a two-prisms camera (dispersion 42 A/mm at ${\rm H}_{2}$), do not leave any doubt about the absence of any emission line, even in the red region. ${\rm H}_{\alpha}$ itself appears strong in absorption.

Astrophysical Observatory, Asiago April 8, 1966

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References

Andrews A.D., Armagh Contribution No. 46, 1964 Andrews A.D., Armagh Contribution No. 49, 1965