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The star probably belongs to the group of the so called nova-like variables or to the U Gem stars. The light-outbursts take place in intervals of about 1 500 days, the limits of brightness being  $12.^m_5$  and  $15.^m_5$  photographically. Superposed on these long term variations are secondary changes with durations less than  $100^d$  and with amplitudes  $1.^m_0$  or more. Night runs of plates show some evidence of still shorter variations (order of magnitude: hours), but these need confirmation by more observations.

Spectral type according to Herbig (ApJ 131, p. 632) G2,  
luminosity class IV or V.

For further details see forthcoming paper in MVS.

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