

COMMISSION 27 OF THE I.A.U.
INFORMATION BULLETIN ON VARIABLE STARS
NUMBER 105

Konkoly Observatory
Budapest
13 September 1965

V 828 CYGNI

On the request of Mr. Meinunger the star was observed photoelectrically on several nights (without filters, multiplier 1 P 21). Following the observations the star is not a semiregular variable with a period of $30^d \pm$ (as listed in the GCVS Supp.I), but a β Lyraestar with the elements

$$\text{Min}_{\odot} = 243\ 8937.476 + 2^d.10454\ E.$$

Both minima are of equal depth; the range of variation is $0^m.42$. The spectral type according to Herbig is B8.

The star is recommended for observation being its brightness $9^m.7$ in maximum.

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